Constraining (n,y) reaction cross sections for astrophysical applications

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Overview

R-process nucleosynthesis
 Uncertainties

 Neutron capture rates

• Experiment

- Results
- Future plans



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Nucleosynthesis paths





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Open questions: What is the site of the r-process?



Credit: Erin O'Donnell, MSU

Core Collapse Supernova?

Neutron Star Merger?







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r-process calculations



• Abundance pattern is different for the different astrophysical scenarios.

- Does one of them reproduce
- the observed abundances best?
- Why can't we tell?



National Science Foundation Michigan State University M. Mumpower, J. Cass, G. Passucci, R. Surman, A. Aprahamian, AIP Adv. 4, 041009 (2014)

r-process





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Nuclear Physics Uncertainties: (n,γ)



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Current (n, γ) measurements





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Neutron Capture – Uncertainties



Traditional Oslo method

- Reaction based
- Applicable close to stability
- Populate the compound nucleus of interest through a transfer or inelastic scattering reaction
- \succ Extract level density and γ -ray strength function
- >Calculate "semi-experimental" (n, γ) cross section
- > Excellent agreement with measured (n, γ) reaction cross section





- Populate the compound nucleus via β-decay (large Q-value far from stability)
- Spin selectivity correct for it
- \bullet Extract level density and $\gamma\text{-}ray$ strength function
- Advantage: Can reach (n,γ) reactions where beam intensity is 1 pps.



Spyrou, Liddick, Larsen, Guttormsen, et al, PRL2014

National Superconducting Cyclotron Lab





Summing NaI - SuN



$$E_x = E_{\gamma 1} + E_{\gamma 2} + E_{\gamma 3} + E_{\gamma 4} + \dots$$



16x16 inch
45 mm borehole
2 pieces
8 segments
24 PMTs
Efficiency > 85% for 1 MeV

A. Simon, S.J. Quinn, A.S., et al., Nucl. Instr. Meth A 703, 16 (2013)



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Proof-of-principle: $^{75}Ge(n,\gamma)^{76}Ge$

z	73Se 7.15 H 8: 100.00%	74Se STABLE 0.89%	75Se 119.79 D 8: 100.00%	76Se STABLE 9.37%	77Se STABLE 7.63%	785e STABLE 23.77%	79Se 2.95E+5 Υ β-: 100.00%	80Se STABLE 49.61% 2β-	81Se 18.45 Μ β-: 100.00%
33	72As 26.0 H 8: 100.00%	73As 80.30 D 8: 100.00%5	74As 17.77 D ε: 66.00% β-: 34.00%	75A2 STABLE 10075	76A≗ 1.0942 D β-: 100.00%	77A≗ 38.83 H β-: 100.00%5	78As 90.7 M β-: 100.00%	79As 9.01 M β-: 100.00%	80Αs 15.2 S β-: 100.00%
32	71Ge 11.43 D 8: 100.00%	72Ge STABLE 27.45%	730e STABLE 7.75%	74Ge STABLE 36.50%	75Ge 82.78 M β-: 100.00%	76Ge STABLE 1.73%	77Ge 11.30 H β-: 100.00% β ⁻	78Ge 88.0 Μ β-: 100.00%	79Ge 18.98 S β-: 100.00%
31	70Ga 21.14 M β-: 99.59% 8: 0.41%	71Ga STABLE 39.892%	72Ga 14.10 H β-: 100.00%	73Ca 4.86 H β-: 100.00%	74Ga 8.12 M β-: 100.00%	γ 750a 126 S β-: 100.00%	760a 32.6 S β-: 100.00%	77Ga 13.2 S β-: 100.00%	78Ga 5.09 S β-: 100.00%
30	69Zn 56.4 M β-: 100.00%	70Zn ≥2.3E+17 Y 0.61% 2β-	712n 2.45 M β-: 100.00%	72Zn 46.5 H β-: 100.00%	73Zn 23.5 S β-: 100.00%	742n 95.6 S β-: 100.00%	752n 10.2 S β-: 100.00%	762n 5.7 S β-: 100.00%	77Zn 2.08 S β-: 100.00%
	39	40	41	42	43	44	45	46	N

⁷⁶Ga:
$$T_{1/2} = 32.6 \text{ s}$$

 $Q_{\beta} = 7.0 \text{ MeV}$
 $S_n (^{76}\text{Ge}) = 9.4 \text{ MeV}$



National Science Foundation Michigan State University Spyrou, Liddick, Larsen, Guttormsen, et al, PRL2014

Proof-of-principle: $^{75}Ge(n,\gamma)^{76}Ge$





National Science Foundation Michigan State University Spyrou, Liddick, Larsen, Guttormsen, et al, PRL2014

Proof-of-principle: $^{75}Ge(n,\gamma)^{76}Ge$





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Results: ${}^{75}Ge(n,\gamma){}^{76}Ge$





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Weak r-process measurements





Summary and Applicability



- Delayed neutron emission



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Collaboration

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