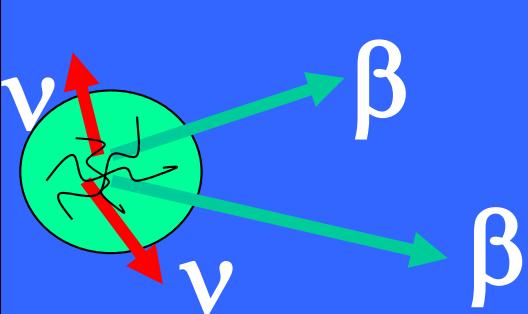
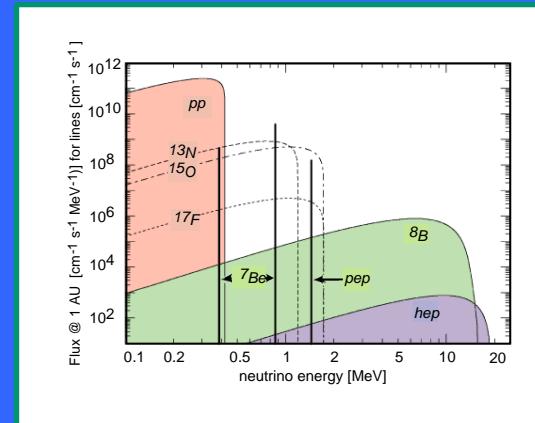


## Charge-exchange reactions GT-transitions, $\beta\beta$ -decay



and  
things beyond



# Outline

➤ Chargex-reactions ( $^3\text{He}, \text{t}$ ) & ( $\text{d}, ^2\text{He}$ )

➤ highlights & features of  $2\nu\beta\beta$  nuclear matrix elements (NME)

$^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$ ,  $^{136}\text{Xe}$

fragmentation – smallest/largest NME



➤ the  $0\nu\beta\beta$  decay nuclear matrix elements

1<sup>st</sup> forbidden NME's and 2- states

➤ solar  $\nu$  SNU rates and ( $^3\text{He}, \text{t}$ ) reaction

$^{71}\text{Ga}(^3\text{He}, \text{t})$ ,  $^{82}\text{Se}(^3\text{He}, \text{t})$

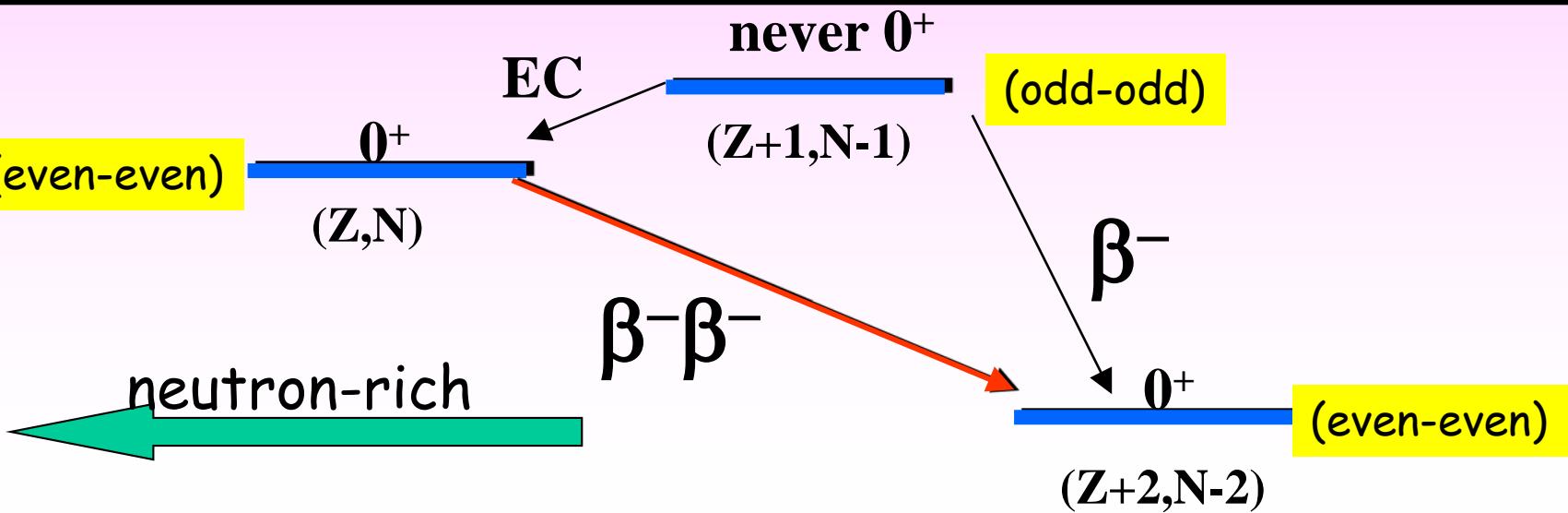
➤ the A=96 system

30 min

the  $^{96}\text{Zr} (\beta^-) \rightarrow ^{96}\text{Nb}$  Q-value  
and a direct test of  $0\nu\beta\beta$  NME



# $\beta^- \beta^-$ decay



$2\nu\beta^- \beta^-$  decay:

$$T_{1/2} \approx 10^{19-21} \text{ y}$$

$$\Gamma = (\text{ph-spc})_{\text{5-body}} \times \left| NME_{\text{allowed}} \right|^2$$

$0\nu\beta^- \beta^-$  decay:

$$T_{1/2} > 10^{24} \text{ y}$$

$$\Gamma = (\text{ph-spc})_{\text{3-body}} \times \left| NME_{\text{any degree}} \right|^2 \times \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|^2$$

# recall: neutrino mass problem

$$\Gamma \propto \left| NME \right|^2 \cdot \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|^2$$

$$U = V \cdot \text{diag}(e^{-i\Phi_1}, e^{-i\Phi_2}, 1) \quad \leftarrow \text{2 extra Majorana-Phases}$$

$$V_{\alpha i} = \begin{pmatrix} V_{e1} & V_{e2} & V_{e3} \\ V_{\mu 1} & V_{\mu 2} & V_{\mu 3} \\ V_{\tau 1} & V_{\tau 2} & V_{\tau 3} \end{pmatrix} = \begin{pmatrix} c_{12}c_{13} & c_{13}s_{12} & s_{13}e^{-i\delta} \\ -c_{23}s_{12} - c_{12}s_{13}s_{23}e^{-i\delta} & c_{12}c_{23} - s_{12}s_{13}s_{23}e^{-i\delta} & c_{13}s_{23} \\ s_{12}s_{23} - c_{13}c_{23}s_{13}e^{-i\delta} & -c_{12}s_{23} - c_{23}s_{12}s_{13}e^{-i\delta} & c_{13}c_{23} \end{pmatrix}$$

**known quantities:**

$$\Theta_{12} = 0.6 \pm 0.1 \rightarrow \approx \pi/6$$

$$\Theta_{23} = 0.7 \pm 0.2 \rightarrow \approx \pi/4$$

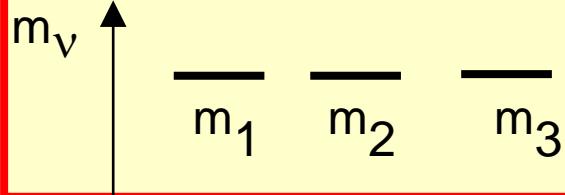
$$\Theta_{13} = 0.11$$

$$\Delta m_{atm}^2 = \left| m_3^2 - m_2^2 \right| \approx 2.6 \times 10^{-3} \text{ eV}^2 \approx \underline{(0.05 \text{ eV})^2}$$

$$\Delta m_{sol}^2 = \left| m_2^2 - m_1^2 \right| \approx 7.9 \times 10^{-5} \text{ eV}^2 \approx \underline{(0.009 \text{ eV})^2}$$

# neutrino-mass-scenarios:

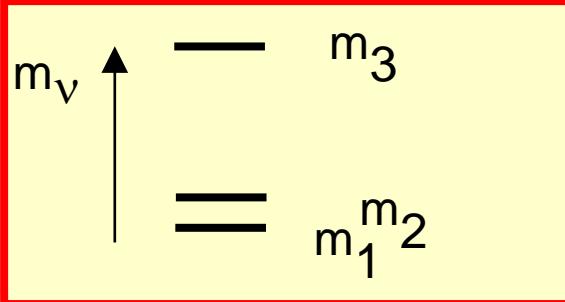
1) degenerate:



$$|m_{\nu_e}| \approx 0.2 \text{ eV}$$

the best of all cases

2) normal hierarchy:

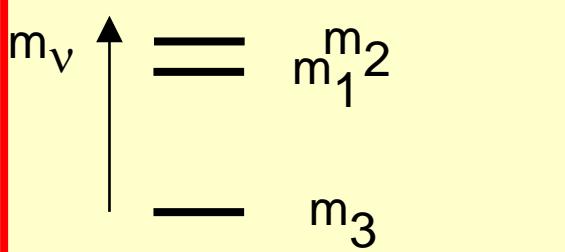


$$|m_{\nu_e}|^2 \propto \Delta m_{sol}^2 \times \left| \frac{3m_1}{\Delta m_{sol}} + e^{-2i(\Phi_2 - \Phi_1)} + (< 0.5)e^{-2i(\delta - \Phi_1)} \right|^2$$

= ZERO!! for:

$$\Theta_{13} \approx 9^\circ \quad (\Phi_2 - \Phi_1) = \frac{\pi}{2} \quad \frac{3m_1}{\Delta m_{sol}} = 1$$

3) inverted hierarchy:



$$|m_{\nu_e}|^2 \propto \Delta m_{atm}^2 \times |3 + e^{-2i(\Phi_2 - \Phi_1)}|^2$$

if inverted hierarchy could be established  
(LHC, SN- $\nu$ , precision-oscillation)

THEN:

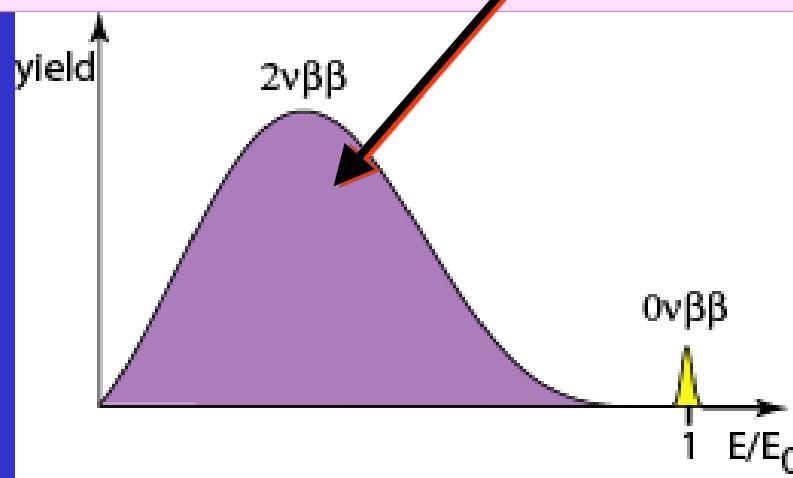
$$|m_{\nu_e}| \approx \Delta m_{atm}$$

or neutrino is a Dirac-particle

NME important

# Nucl. Matrix Elements

## $2\nu\beta^- - \beta^-$ decay



q-transfer like in ordinary  
 $\beta$ -decay

( $q \sim 0.01 \text{ fm}^{-1} \sim 2 \text{ MeV/c}$ )

i.e. only allowed transitions possible

$$G_{(b^- b^-)}^{2n} = \frac{C}{8p^7} \frac{\bar{c}_{Ge} G_F g_A}{\sqrt{2}} \cos(Q_C) \frac{\dot{O}^4}{F} \left| M_{DGT}^{(2n)} \right|^2 F_{(-)}^2 f(Q)$$

$= G^{2n}(Q, Z)$

$\propto Q^{11} \cdot Z^2$

$\left| M_{DGT}^{(2n)} \right|^2$

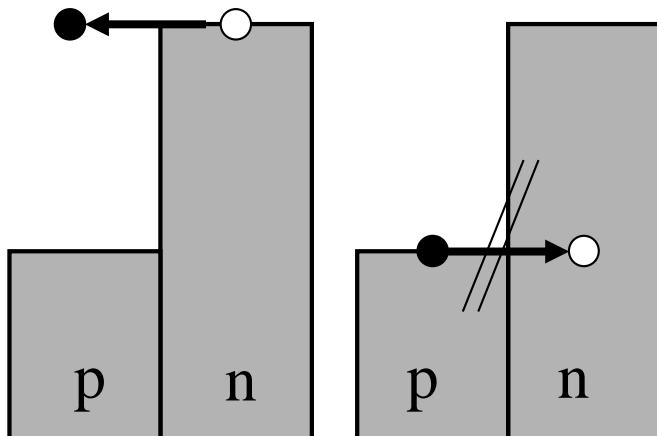
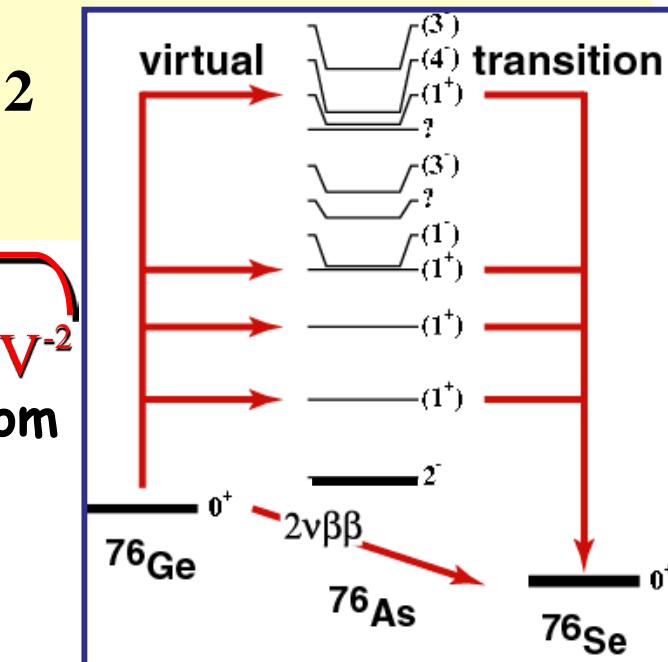
$\exp \approx 10^{-3} \text{ MeV}^{-2}$   
extracted from half-life

**favorable:**

1. high Q-value
2. large Z

**unfavorable (but cannot be changed):**

1. large neutron excess  
(Pauli-blocking)



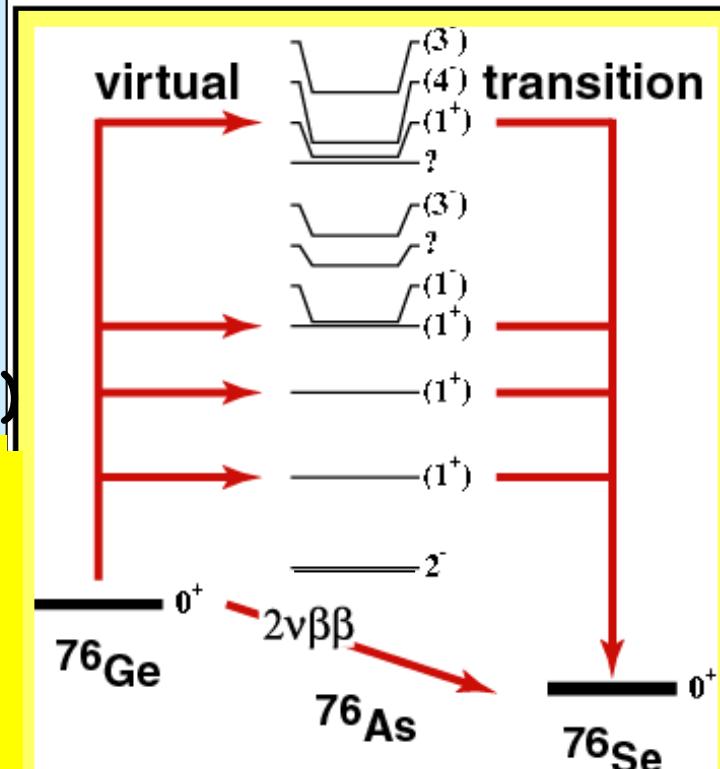
$$M_{\text{DGT}}^{(2n)} = \frac{\langle 0_{g.s.}^{(f)} | \hat{\mathbf{a}}_k s_k t_k | 1_m^+ \rangle \langle 1_m^+ | \hat{\mathbf{a}}_k s_k t_k | 0_{g.s.}^{(i)} \rangle}{\frac{1}{2} Q_{bb}(0_{g.s.}^{(f)}) + E(1_m^+) - E_0}$$

$$= \hat{\mathbf{a}}_m \frac{M_m(GT^+) M_m(GT^-)}{E_m}$$

to remember:

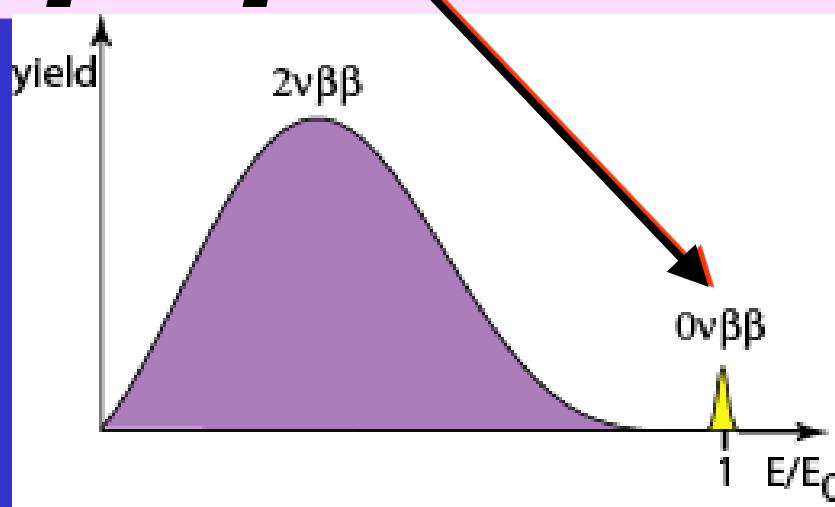
1. 2 sequential & „allowed“  $\beta^-$ -decays of „Gamow-Teller“ type
2. „1, 2, 3, ... forbidden“ decays negligible
3. Fermi-transitions do no contribute (because of different isospin-multiplets)

Can be determined via charge-exchange reactions in the (n,p) and (p,n) direction ( e.g. ( $d, {}^2He$ ) or ( ${}^3He, t$ ) )



# N<sub>ucl.</sub>M<sub>atrix</sub>E<sub>lements</sub>

## $0\nu\beta^-\beta^-$ decay



**neutrino is a virtual particle**  
 $q \sim 0.5 \text{ fm}^{-1}$  ( $\sim 100 \text{ MeV}/c$ )  
(due to Heisenberg  $D q \times D x \sim 1$ )  
degree of forbiddenness is lifted

$G_{(b^- b^-)}^{0n} = G^{0n}(Q, Z) g_A^4 M_{\text{DGT}}^{(0n)} - \frac{c g_V}{c g_A} \frac{\dot{\phi}^2}{F} M_{\text{DF}}^{(0n)} |m_{n_e}|^2$

theory  $\approx 10 !!$   
 largely independent of  $(A, Z)$   
 (except near magic nuclei)

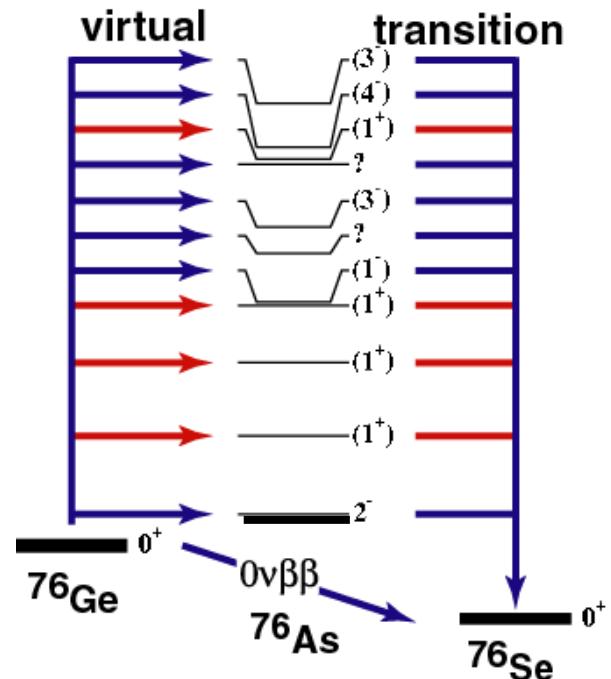
$\propto Q^5 \cdot Z^4$

mass of Majorana- $\nu$  !

to remember:

1. „higher-fold forbidden“ transitions possible
2. Fermi-transitions important
3. „Pauli-blocking“ largely lifted
4. large  $Q$ -value, high  $Z$  important

NOT (easily) accessible via charge-exchange reactions

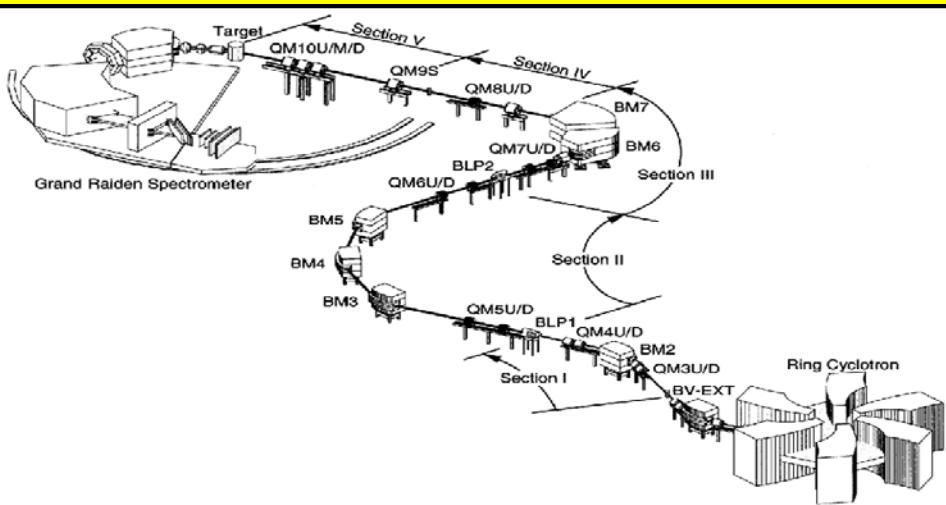


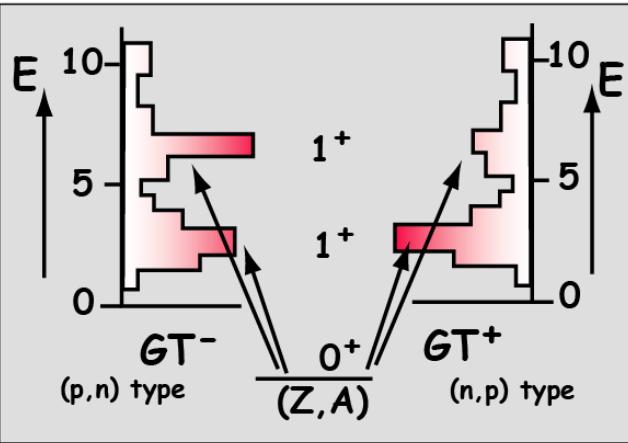
# Charge-exchange reactions

## Grand Raiden Magnetic Spectrometer



$\Delta E/E \sim 5 \times 10^{-5}$  ~ 25 keV  
at 420 MeV ( $^3\text{He}$ )





$$M(GT) = \langle 1^+ || \sigma\tau^+ || 0_{g.s.}^i \rangle$$

$$B(GT) = \frac{1}{2J_i + 1} | M(GT) |^2$$

hadronic probes: ( $n, p$ ), ( $d, {}^2He$ ), ( $t, {}^3He$ )  
or ( $p, n$ ), ( ${}^3He, t$ )

$$\left[ \frac{d\sigma}{d\Omega} \right] = \left[ \frac{\mu}{\pi \hbar} \right]^2 \frac{k_f}{k_i} N_d |v_{\sigma\tau}|^2 | \langle f | \sigma\tau | i \rangle |^2$$

$$q = 0!!$$

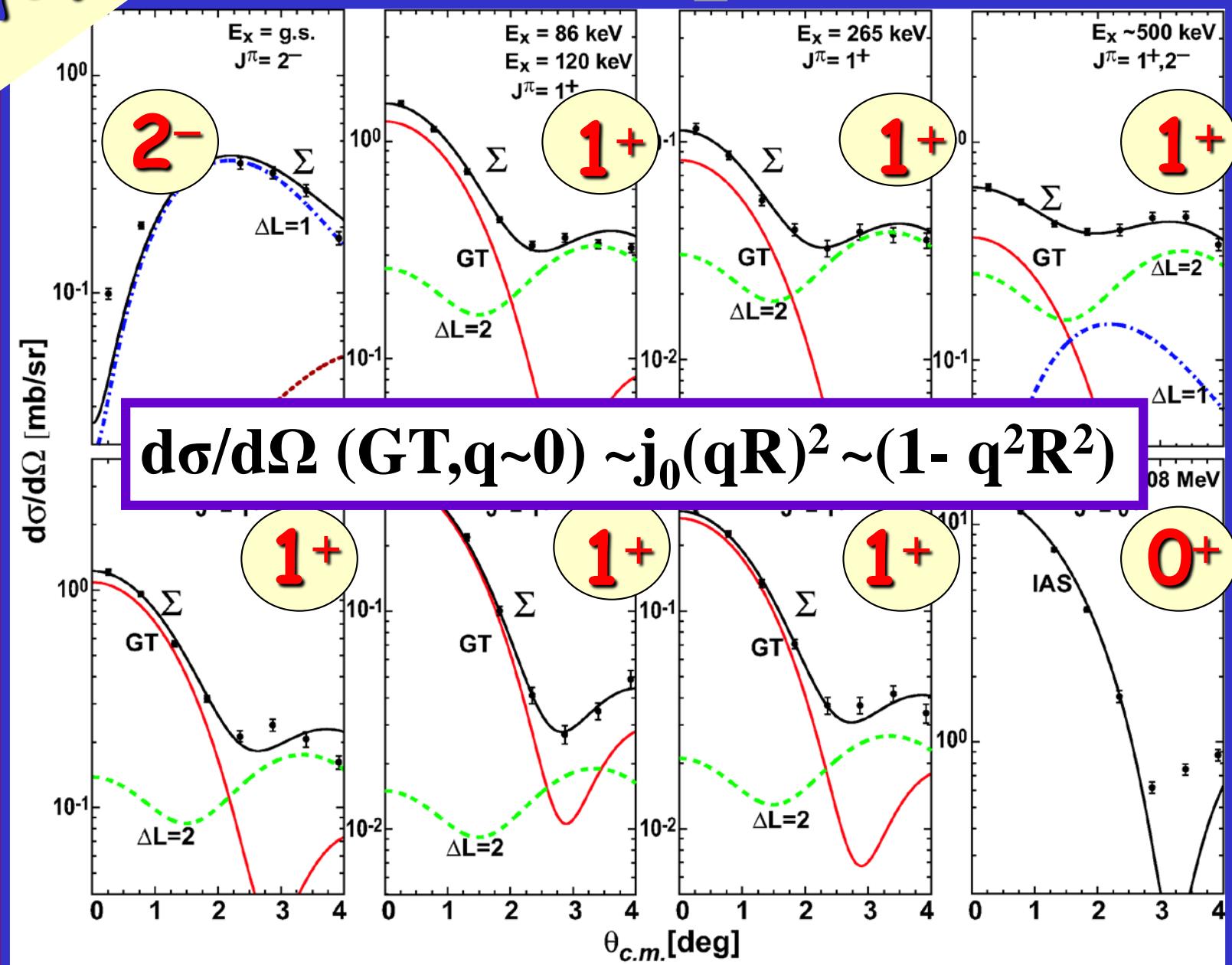
  
largest at 100 - 200 MeV/A

Q: what is the connection  
between „weak  $\sigma\tau$  operator“  
and the hadronic reaction

A: dominance of the  $V\sigma\tau$   
effective interaction at  
medium energies

$^{76}\text{Ge}$ - $^{76}\text{As}$

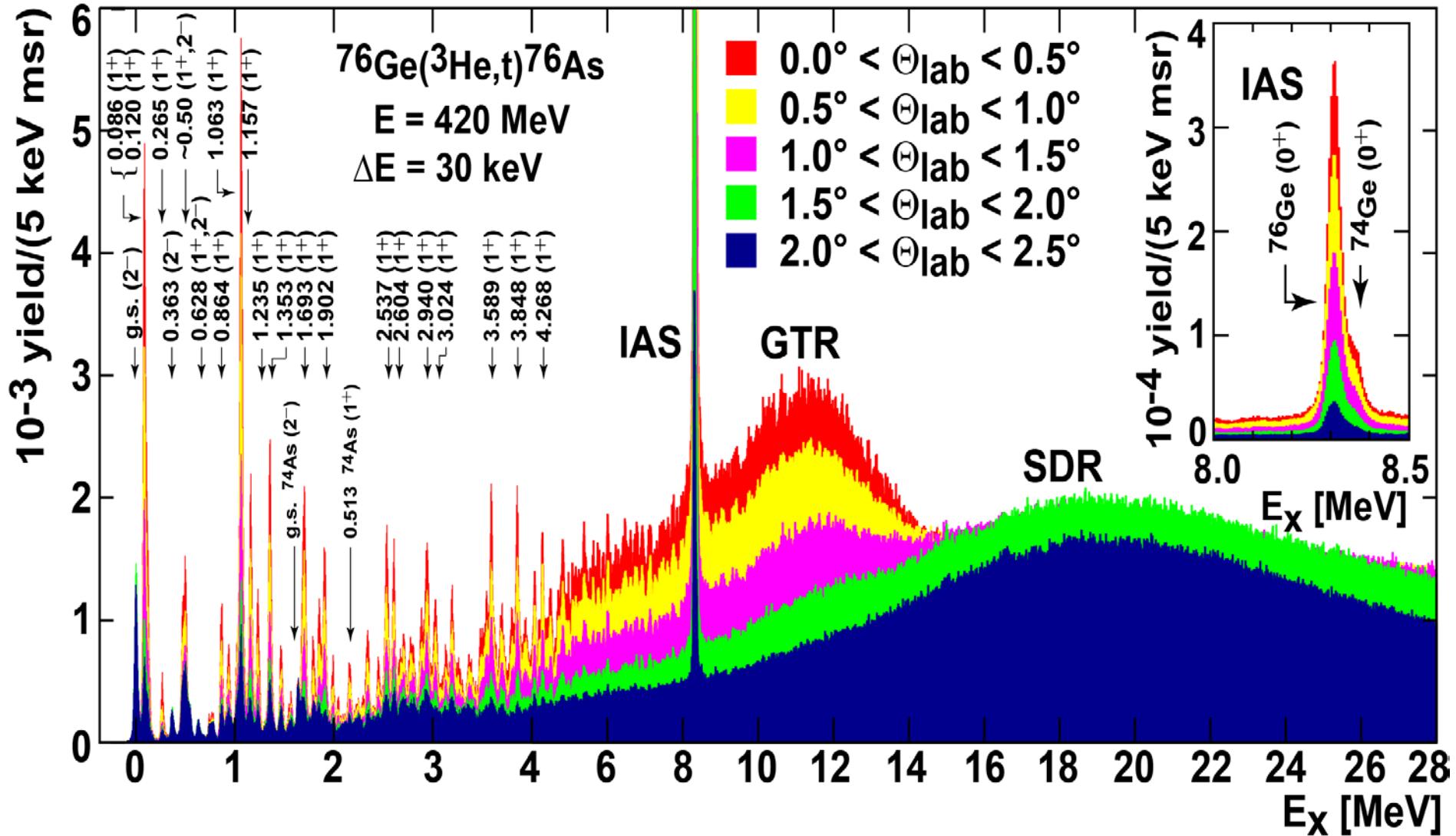
# examples



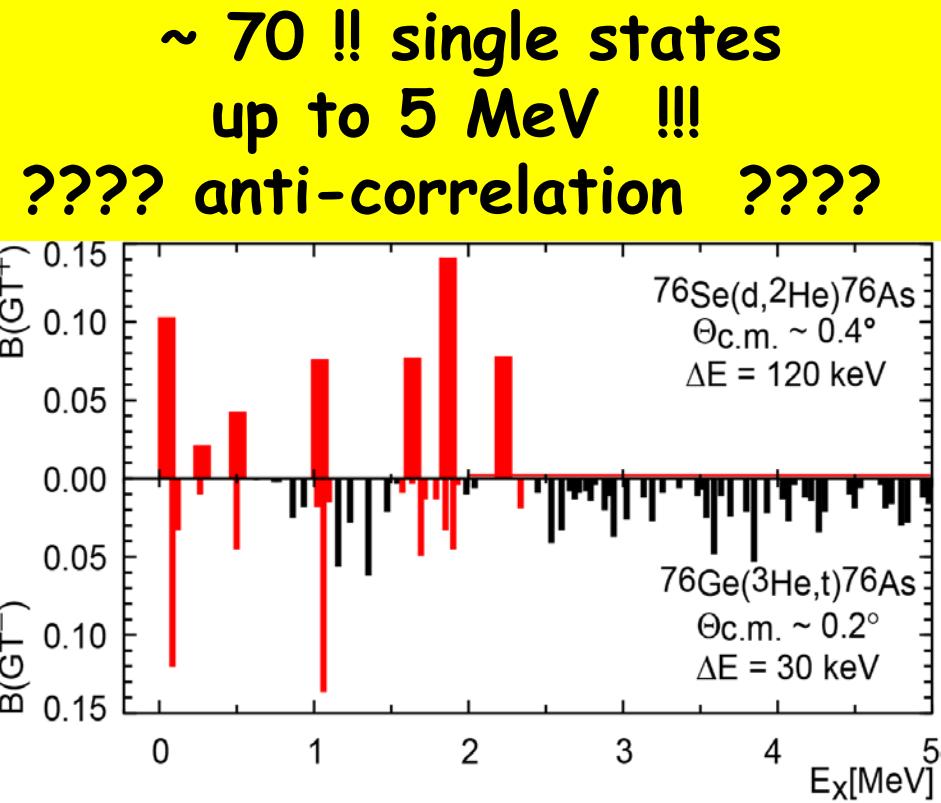
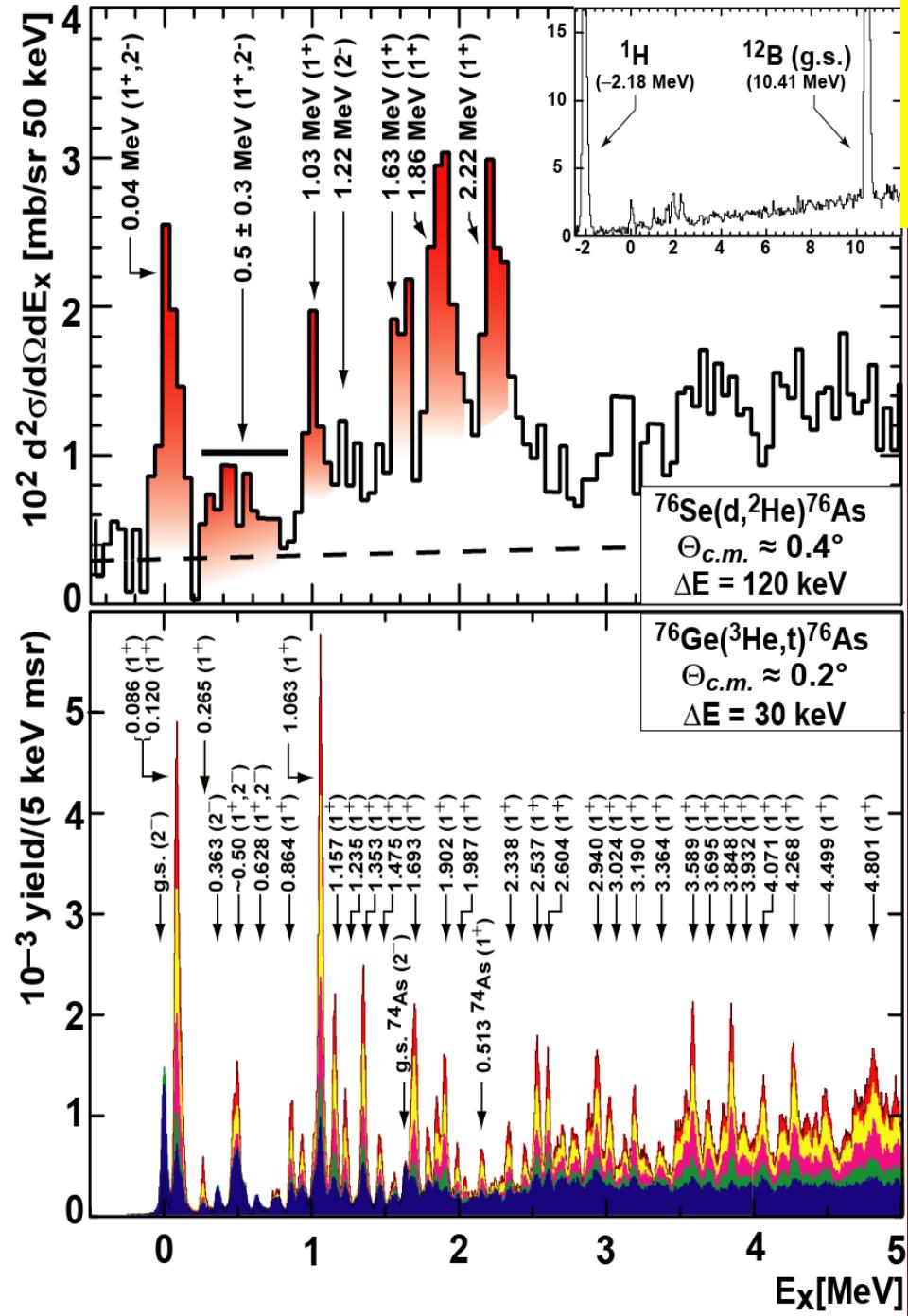
$^{76}\text{Ge}$

$N-Z=10$

**Resolution is the key !!!**



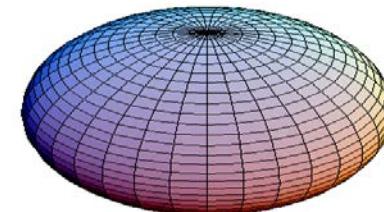
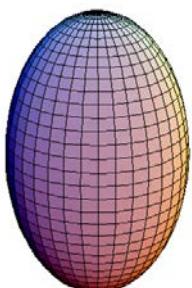
almost 70 !! resolved single states up to 5 MeV  
 identified as GT 1+ transitions !!!



is the anti-correlation a property of deformation ??

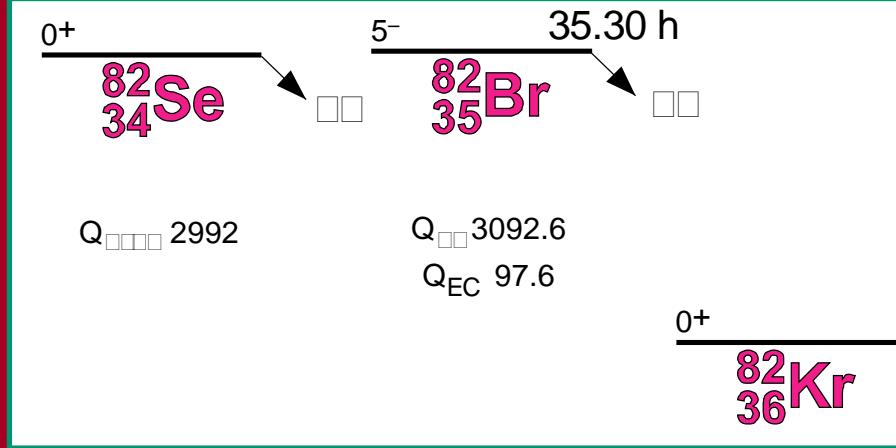
76Ge

76Se



# $^{82}\text{Se}$

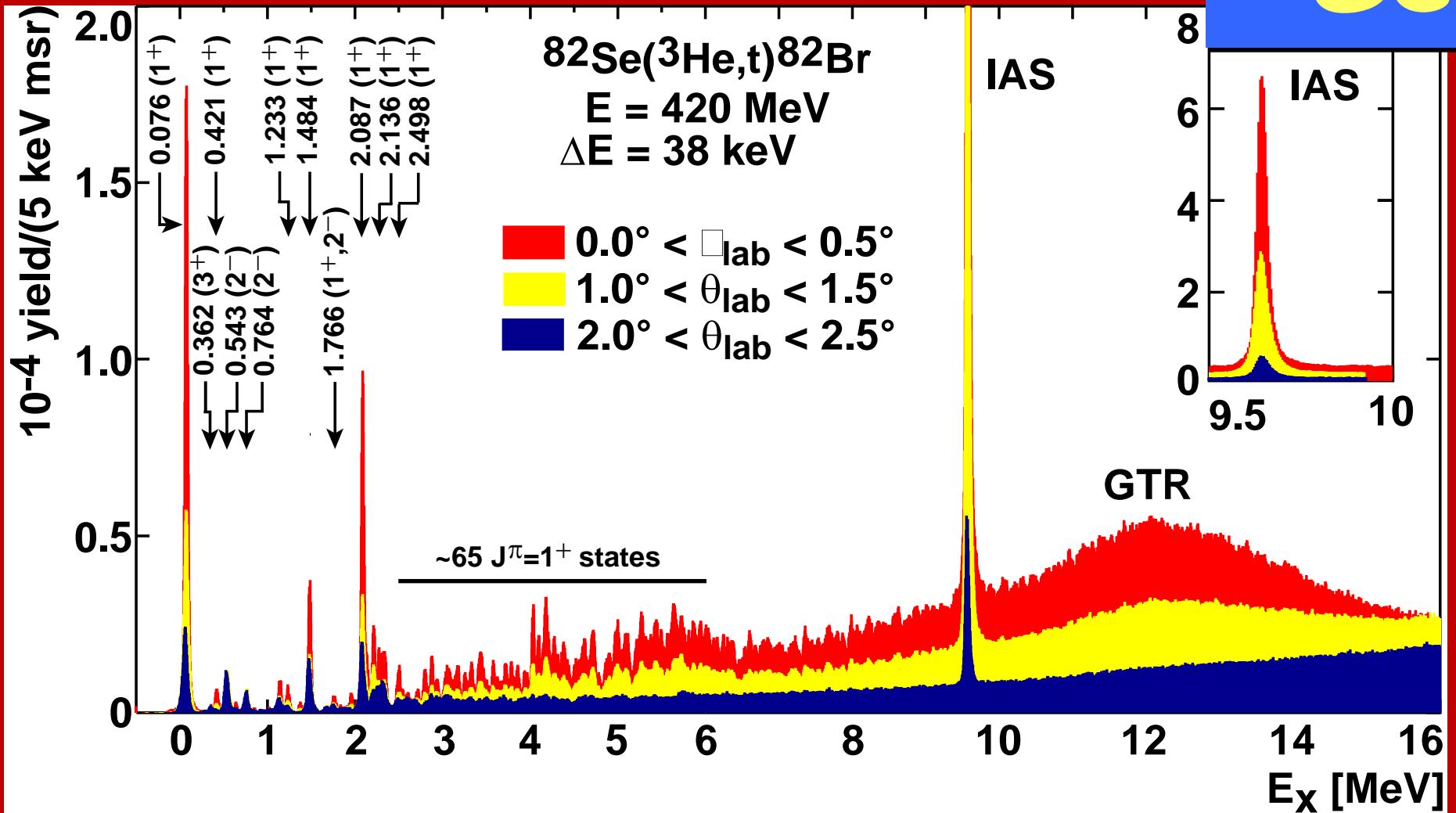
N-Z=14



Resolution is the key !!!

possibly useful for solar neutrino detection

***82*Se**



3 isolated GT transition below 2 MeV-  
fragmentation recedes to GT resonance

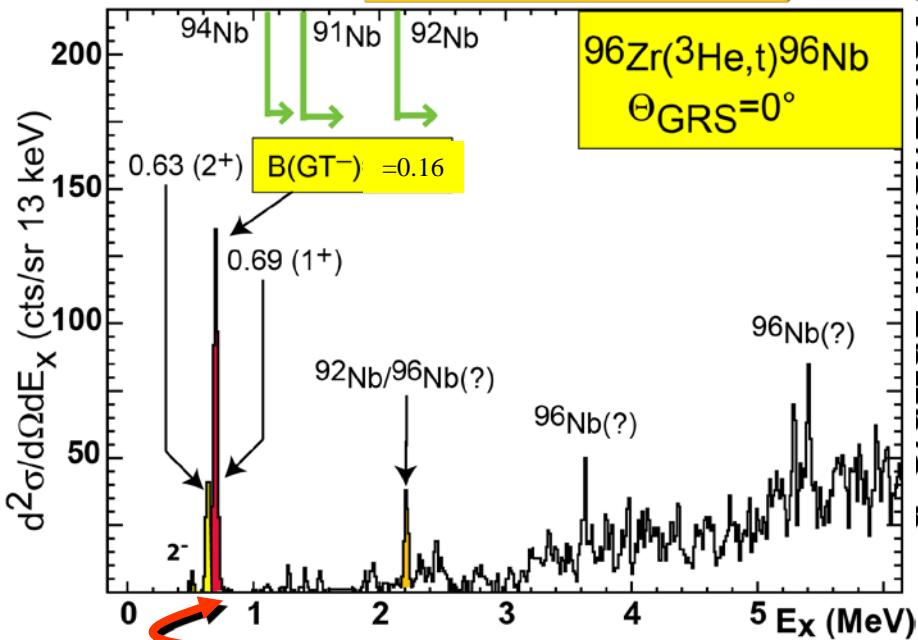
**$^{96}\text{Zr}$**   
 **$N-Z=16$**

**Remember:  $B(\text{GT})_{\text{tot}} = 3(N-Z) \sim 50!$**   
 **$B(\text{F}) = (N-Z)$**

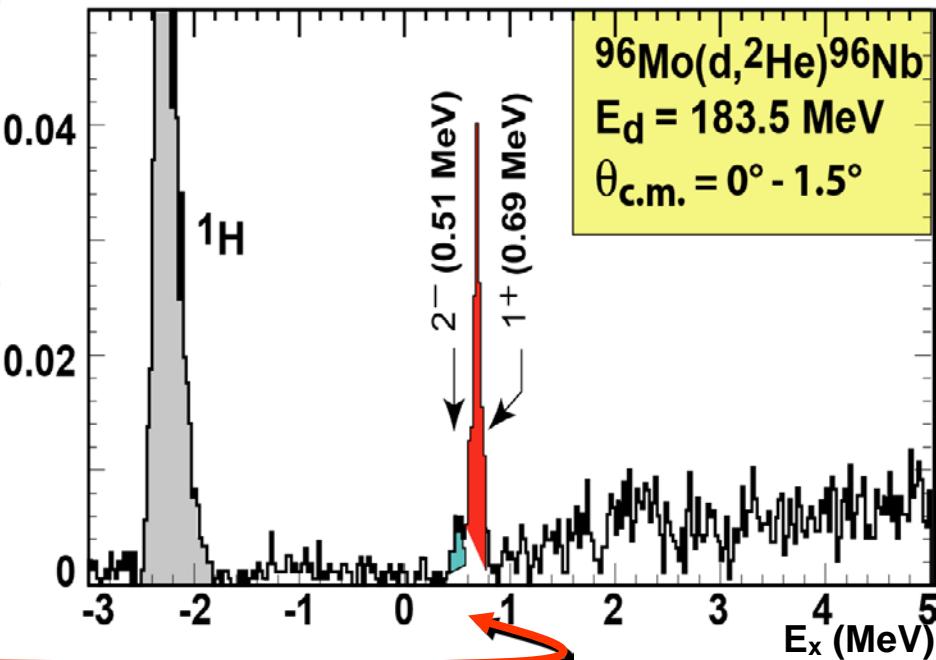
# $(^3\text{He}, \dagger)$

# $(\text{d}, ^2\text{He})$

RCNP 2007/08



$96\text{Zr}(^3\text{He}, \dagger)96\text{Nb}$   
 $\theta_{\text{GRS}} = 0^\circ$



$96\text{Mo}(\text{d}, ^2\text{He})96\text{Nb}$   
 $E_d = 183.5 \text{ MeV}$   
 $\theta_{\text{c.m.}} = 0^\circ - 1.5^\circ$

$$B(GT^-) = 0.16$$

$$B(GT^+) = 0.3$$

Fascination: With only 1 state:

$$T_{1/2}^{\text{calc.}}(2\nu\beta\beta) = (2.1 \pm 0.4) \cdot 10^{19} \text{ years}$$

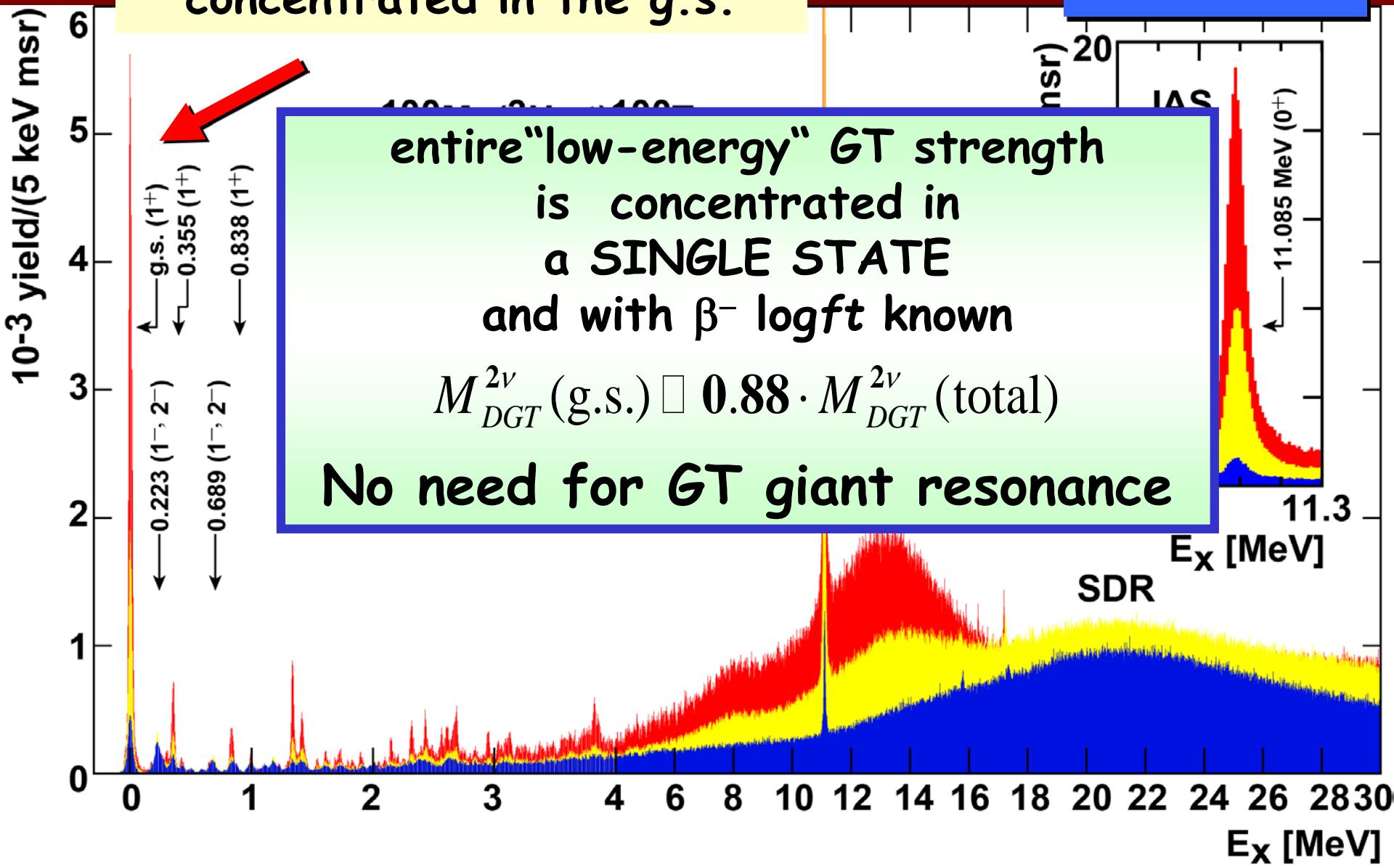
$$T_{1/2}^{\text{exp.}}(2\nu\beta\beta) = (2.3 \pm 0.2) \cdot 10^{19} \text{ years } (\text{NEMO3-result})$$

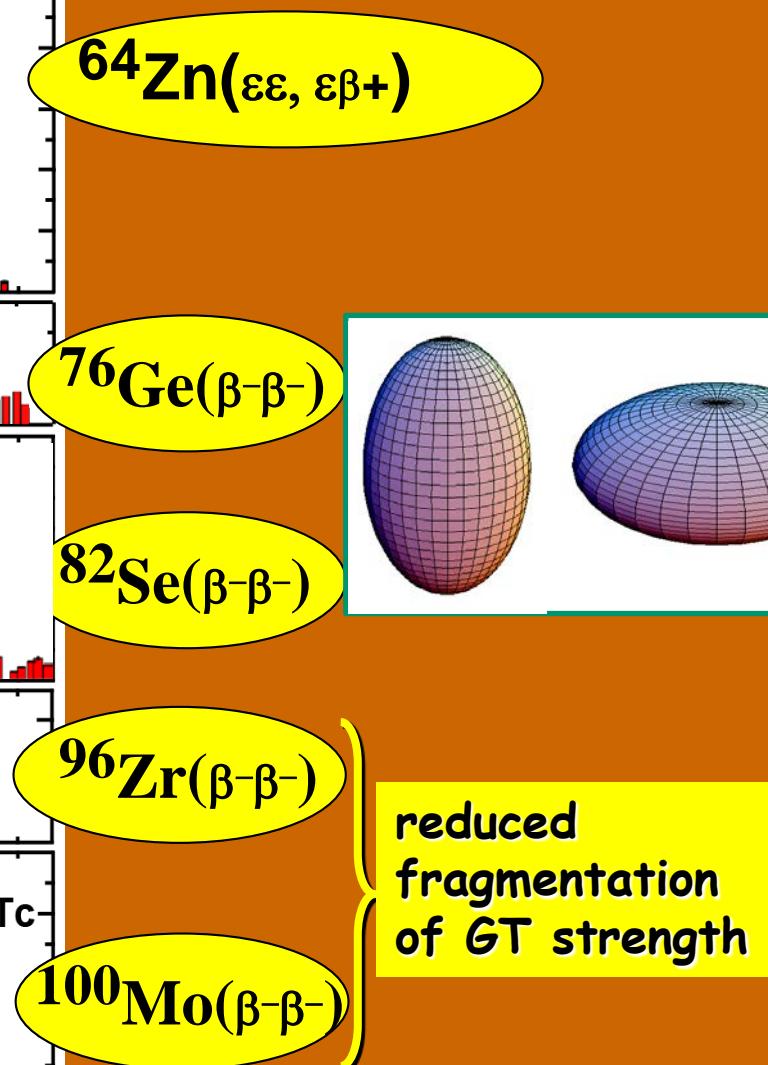
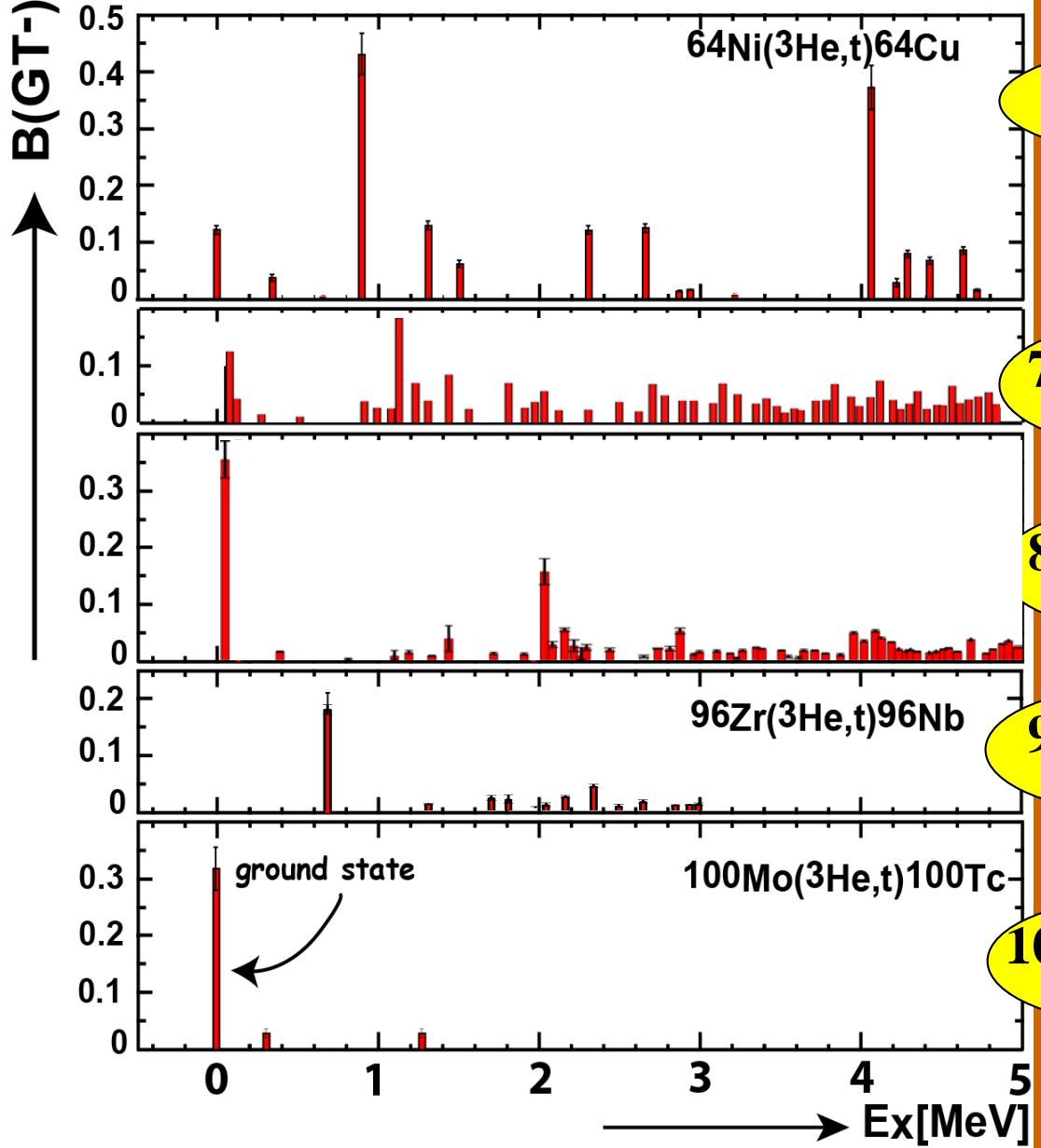
**$^{100}\text{Mo}$**   
 **$N-Z=16$**

**useful as SN neutrino detector  
(sensitive to  $\nu$  temperature in SN)**

HERE: almost the entire low- $E$  GT strength is concentrated in the g.s.

# 100Mo





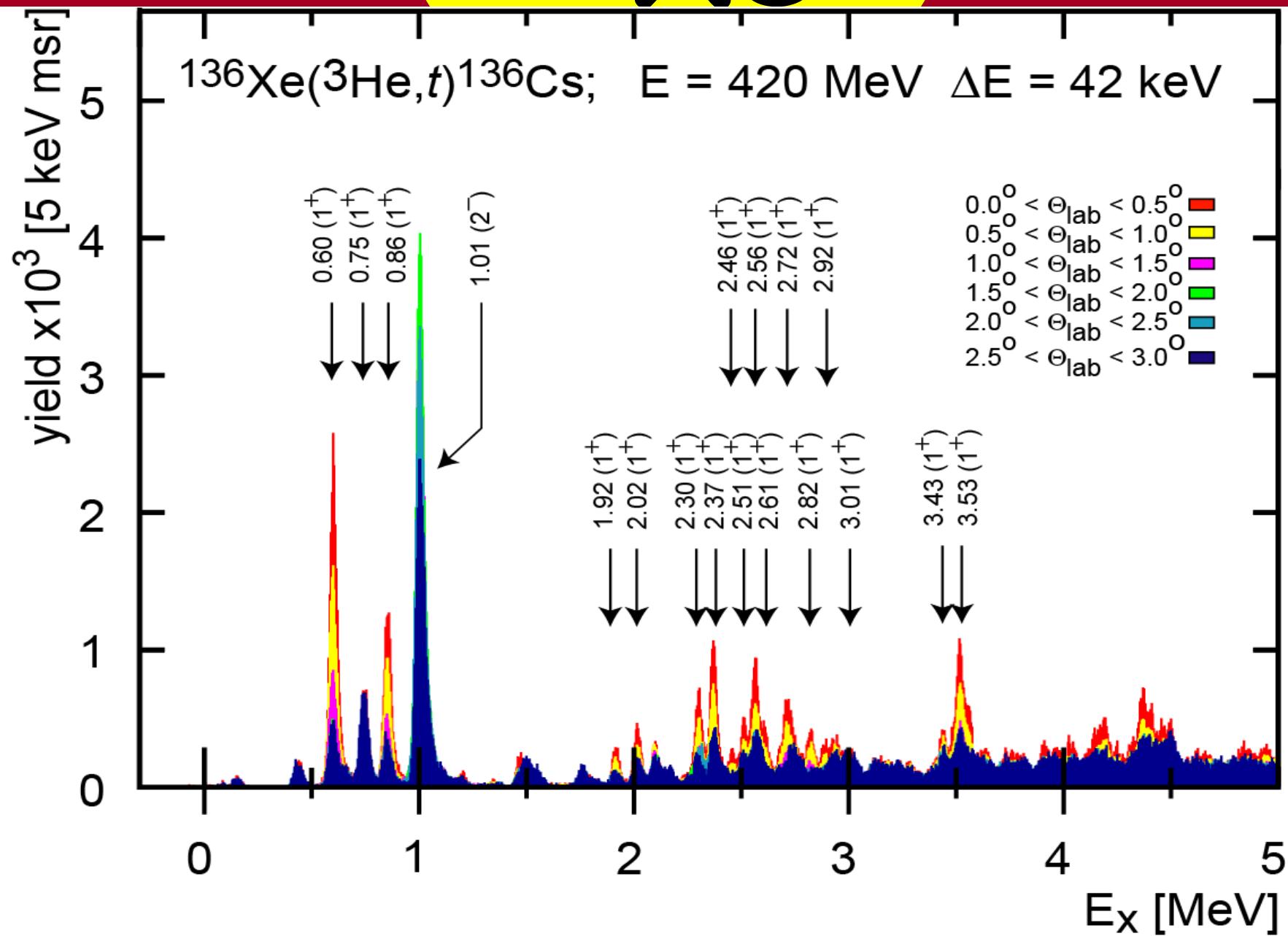
reduced  
fragmentation  
of GT strength

**136Xe**

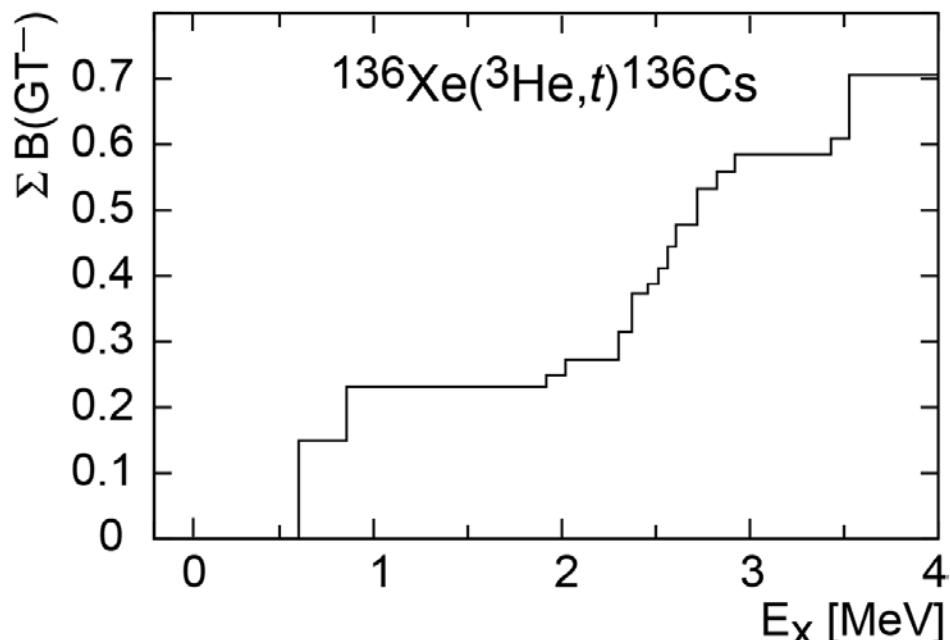
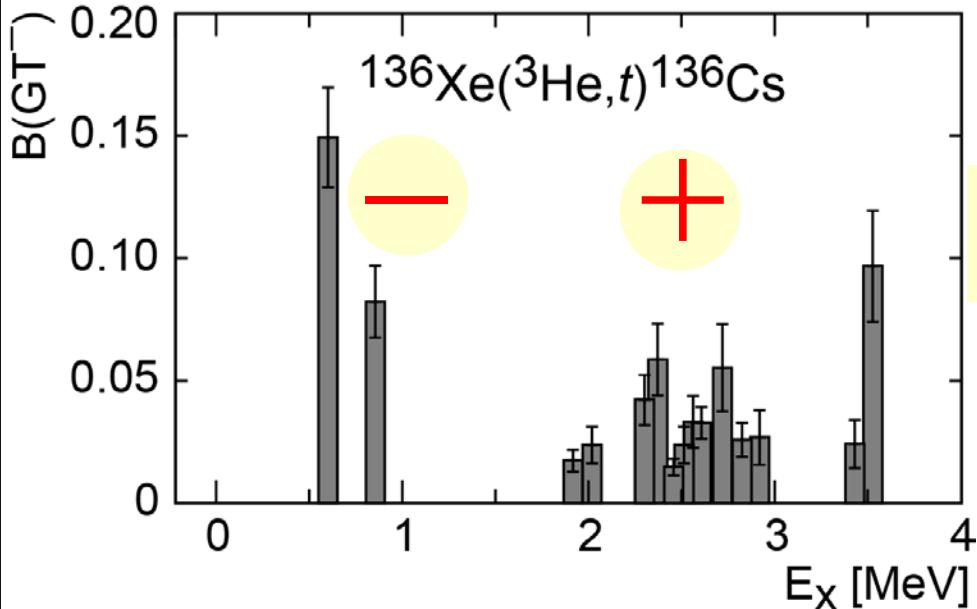
**N-Z=28**

**question: why so stable !!!**

# **136Xe**



# What's the size of the NME?



$$T_{1/2}^{2n} = 2.2 \times 10^{21} \text{ yr}$$

$$M_{\text{DGT}}^{(2n)} \square 0.019 \text{ MeV}^{-1}$$

all signs positive  $\rightarrow$

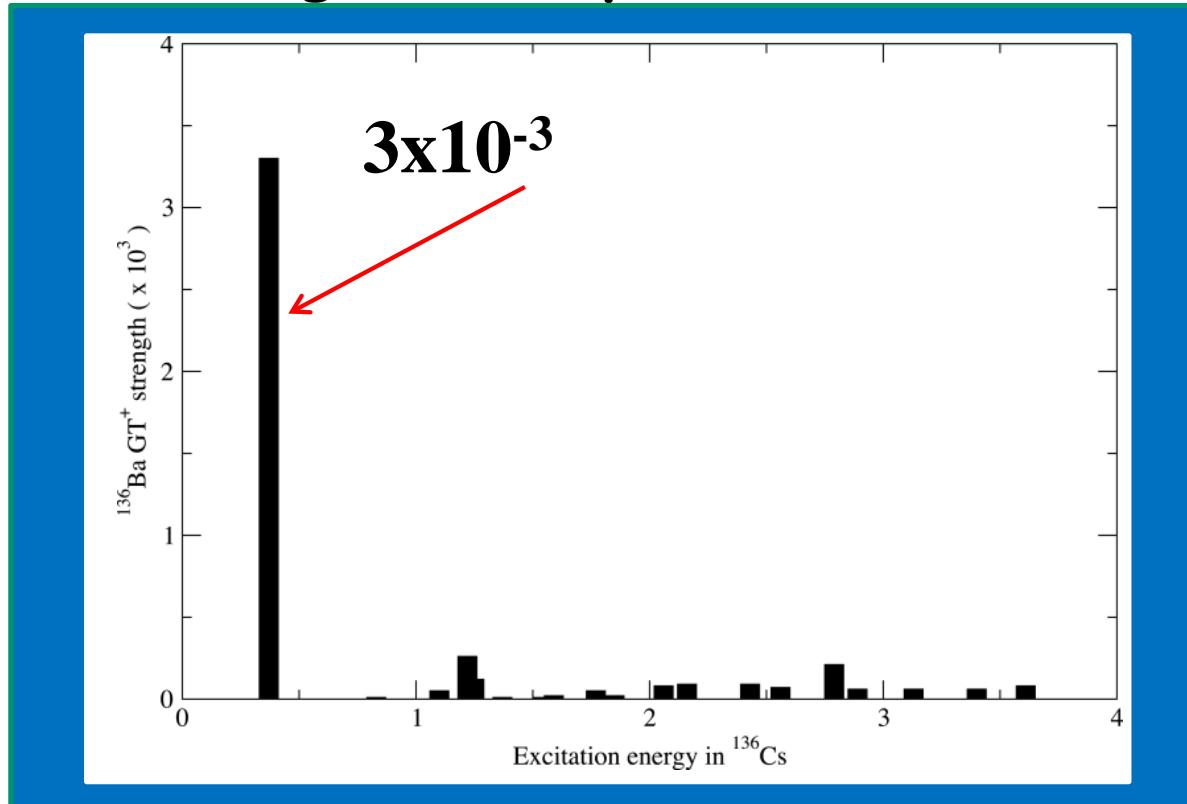
$$B_m(GT^+) \gg 10^{-2} \times B_m(GT^-)$$

$$B_m(GT^+) \gg 10^{-3} !!!!$$

A. Poves (simultaneous to our publication):

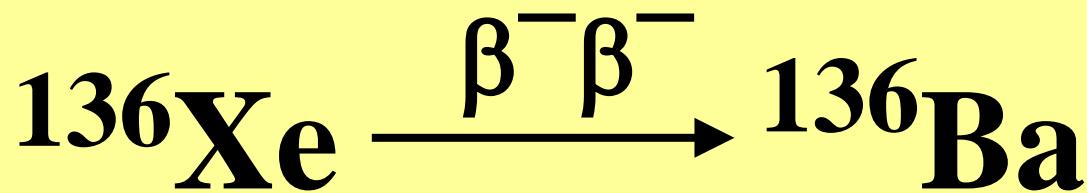
there is no  $B(GT^+)$  strength, except for lowest  $1^+$  state

Recall:  
 $^{136}\text{Xe}$  is almost  
doubly magic!!



Shell model provides conclusive explanation for the  
deemed „pathologically“ long half-life of  $^{136}\text{Xe}$ .

Expt'l test:  $^{136}\text{Ba}(\text{d}, ^2\text{He})^{136}\text{Cs}$

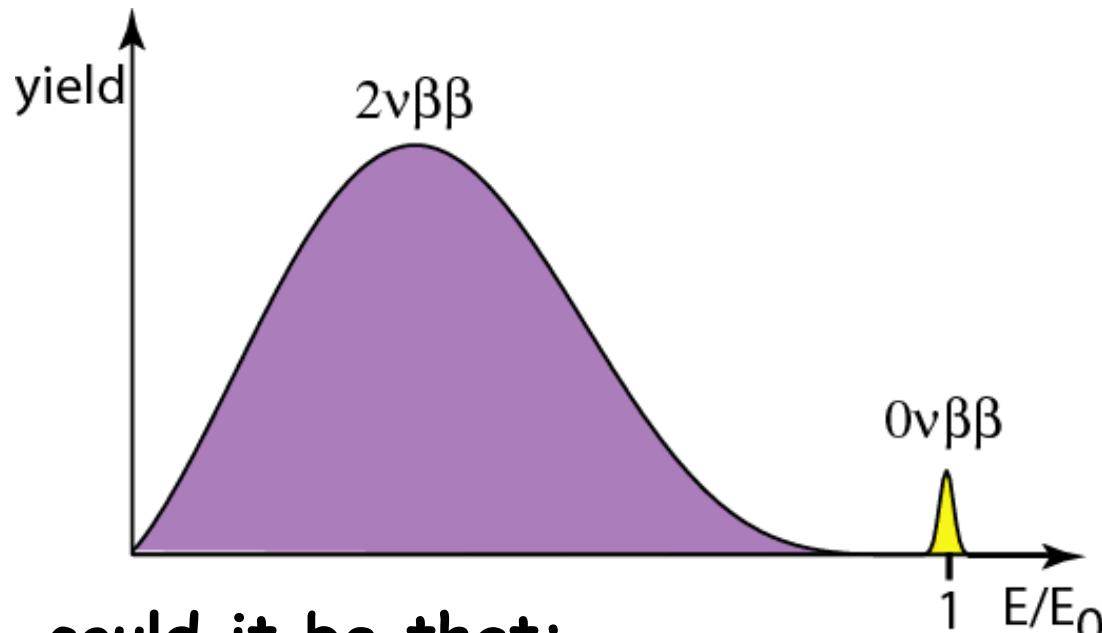


expmt:

question:

$2\nu\beta\beta$  NME is exceptionally small

how does the ME scale in the case of  $0\nu\beta\beta$  decay?



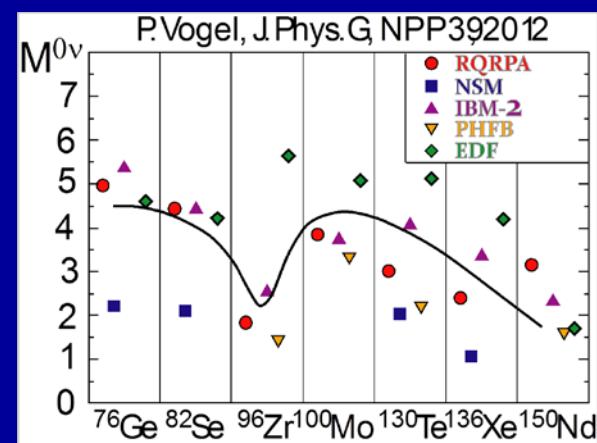
could it be that:

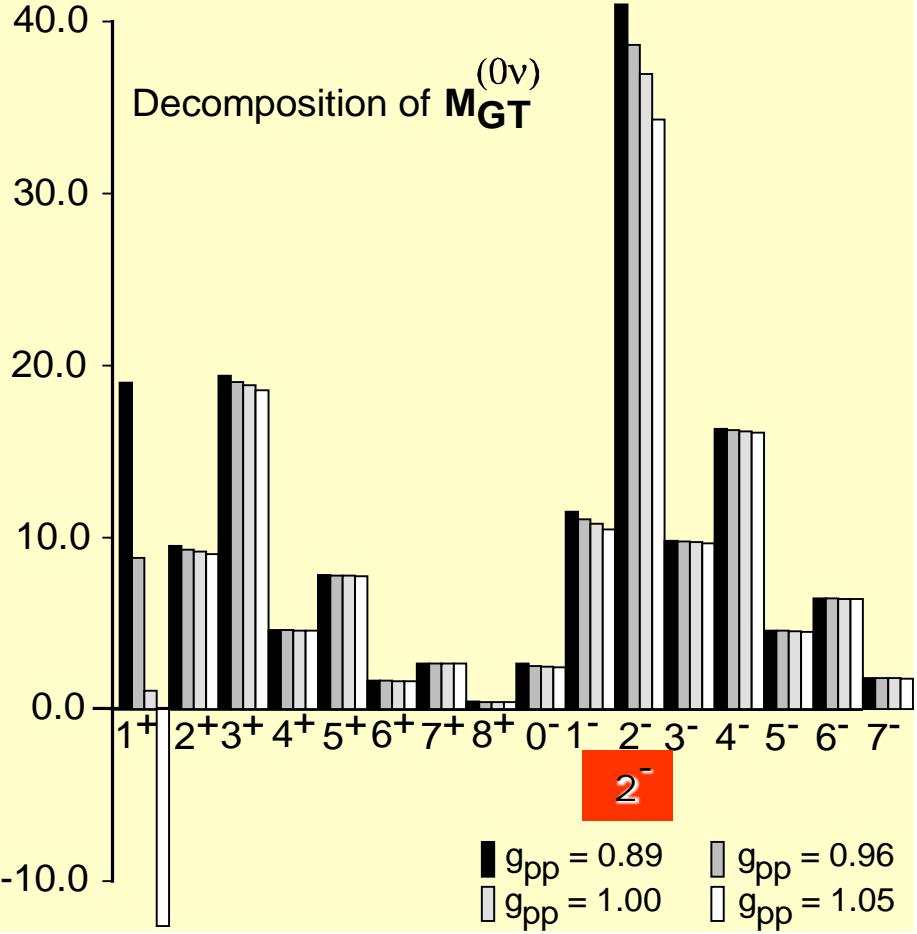
$2\nu\beta\beta$  ME is suppressed **AND**

$0\nu\beta\beta$  ME is enhanced ???

# Experiments towards the $0\nu\beta\beta$ NMEs

Here:  
2<sup>-</sup> states and occupation  
vacancy numbers  
via chargex reactions

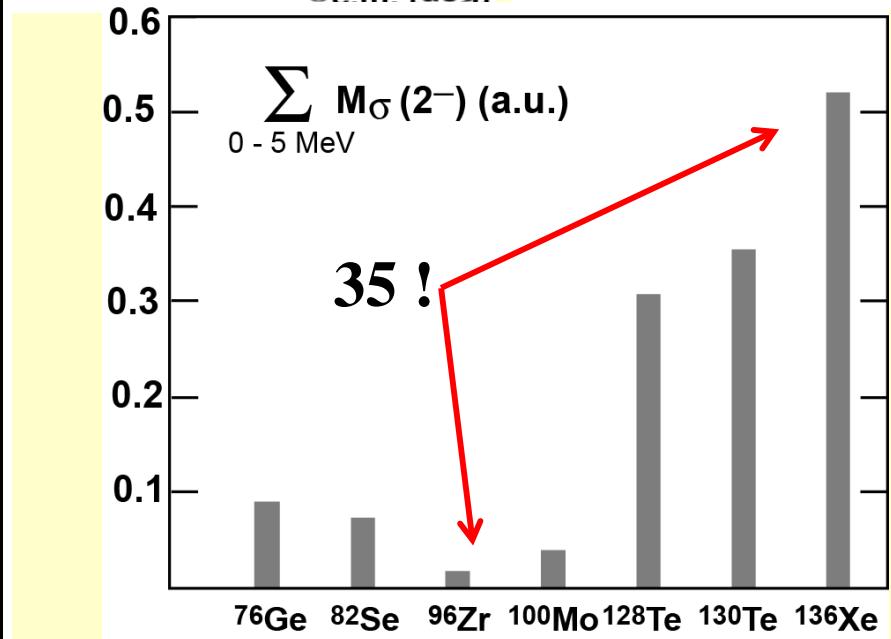
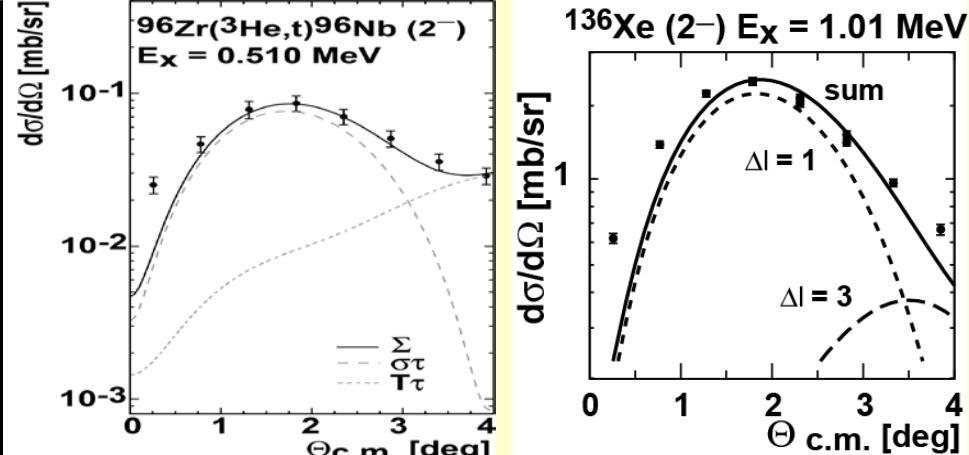




## Theory:

The 2<sup>-</sup> strength makes up  
~ 20-30% of the 0νββ ME!!

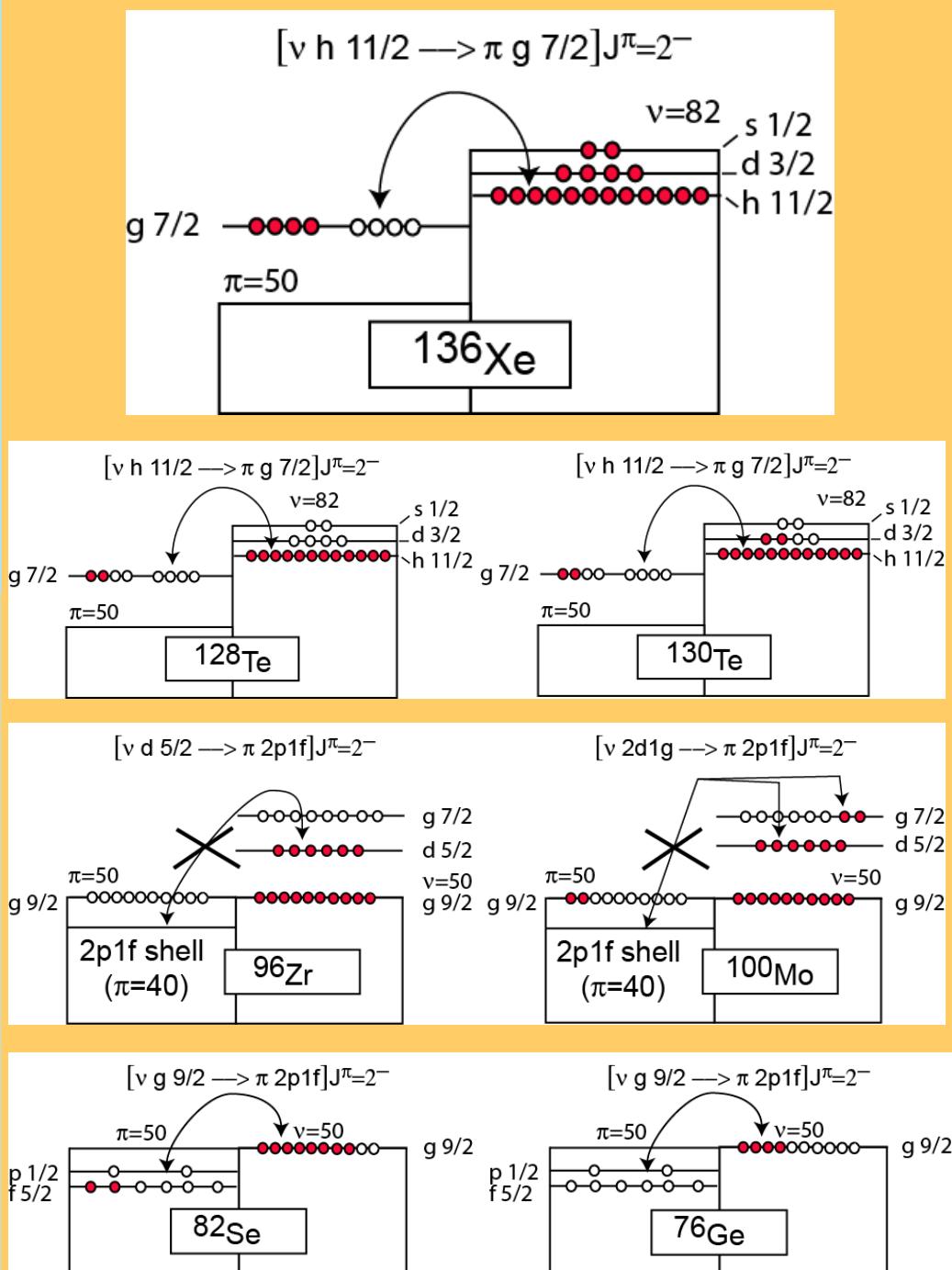
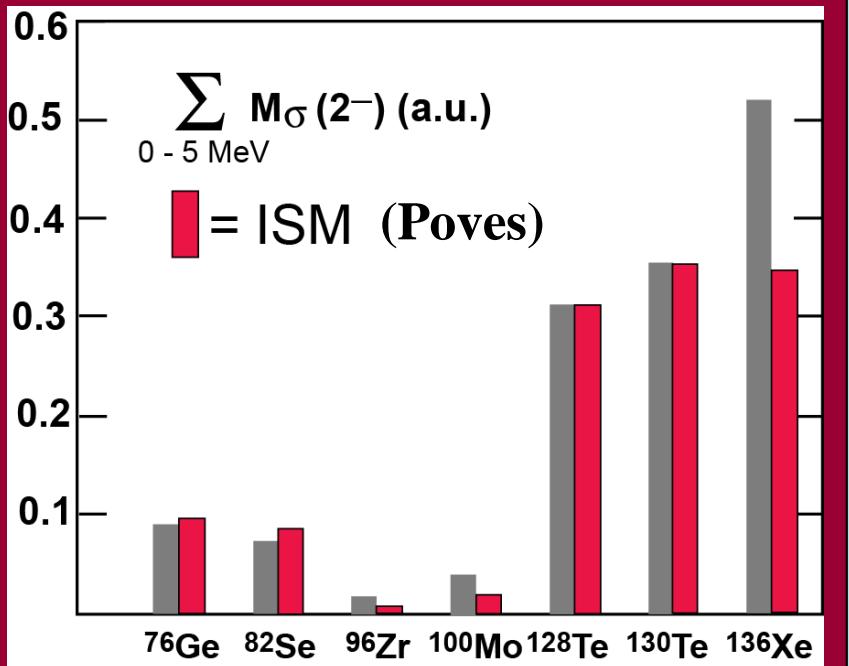
J. Suhonen, Phys. Lett B607, 87 (2005)



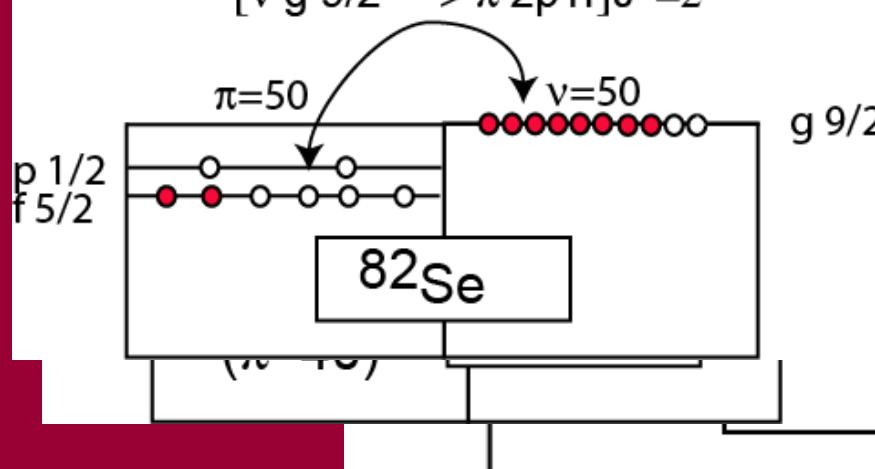
relative 2<sup>-</sup> strength to ~ 5 MeV

## Expt:

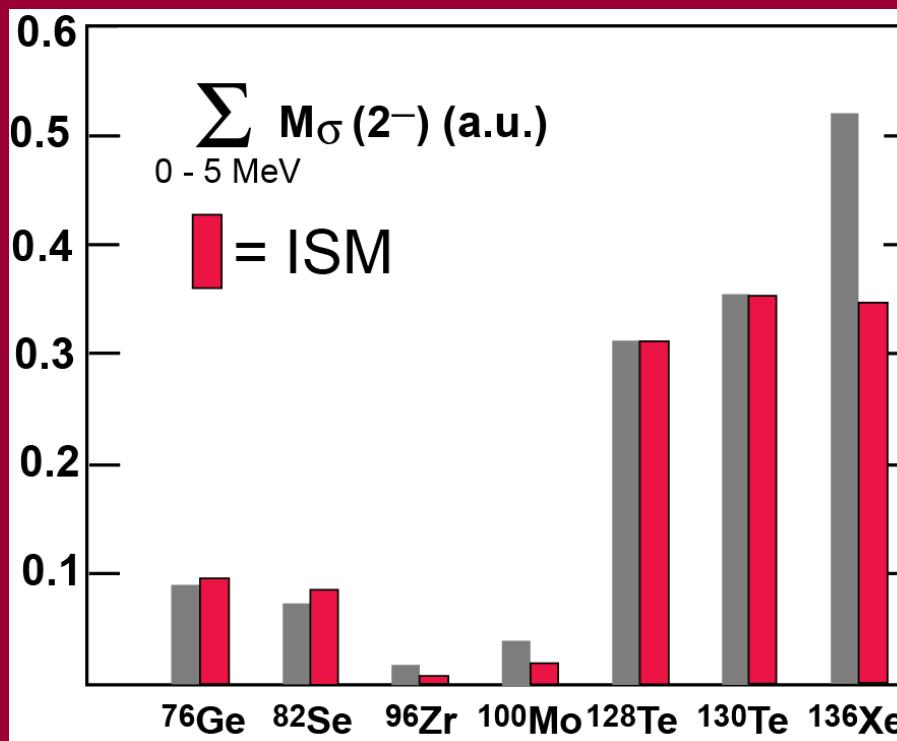
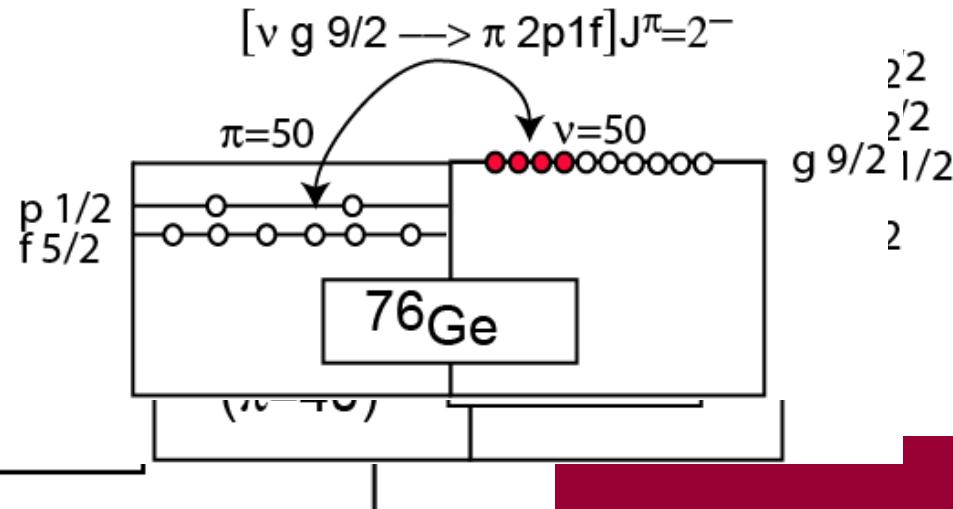
136Xe exhibits largest 2<sup>-</sup> strength  
**0νββ ME enhanced???**

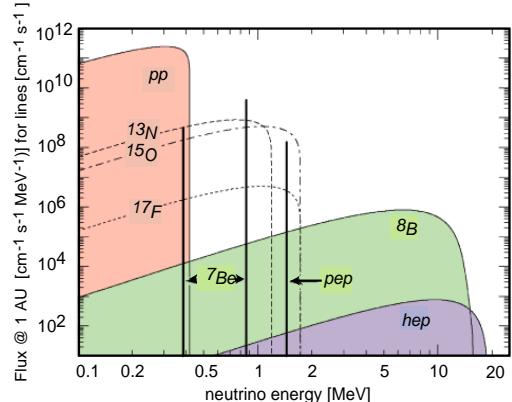


$[\nu \text{ d } 5/2 \rightarrow \pi \text{ 2p1f}] J^\pi = 2^-$



$[\nu \text{ 2d1g} \rightarrow \pi \text{ 2p1f}] J^\pi = 2^-$

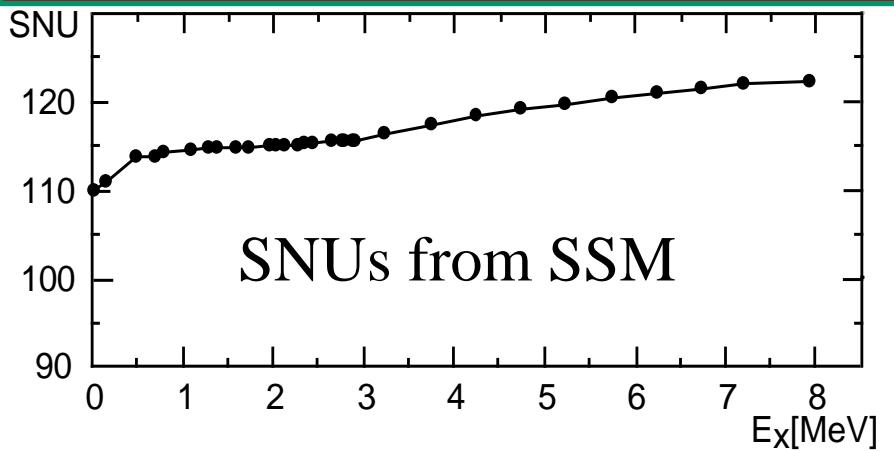
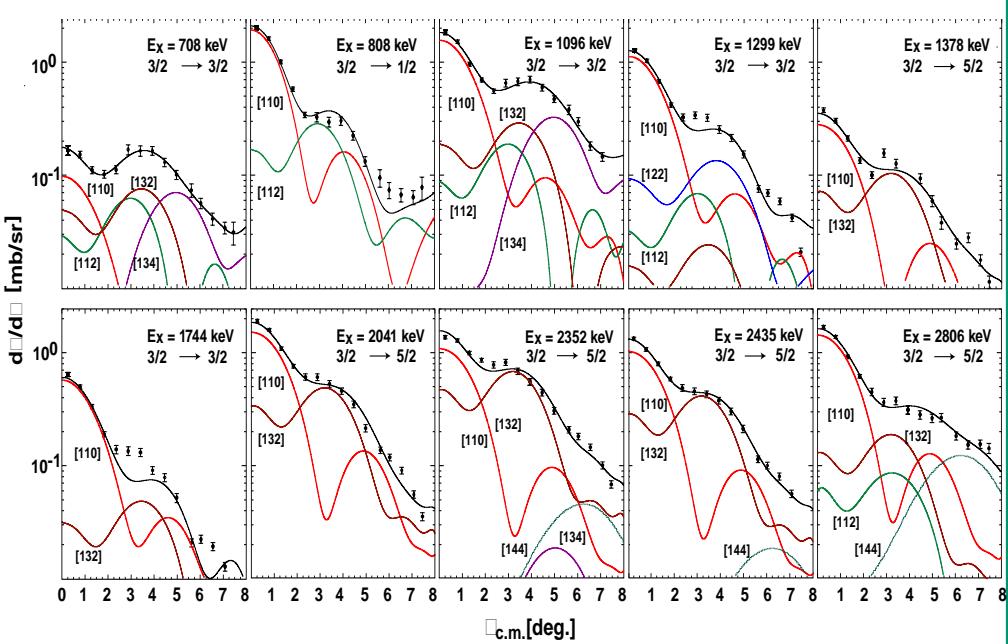
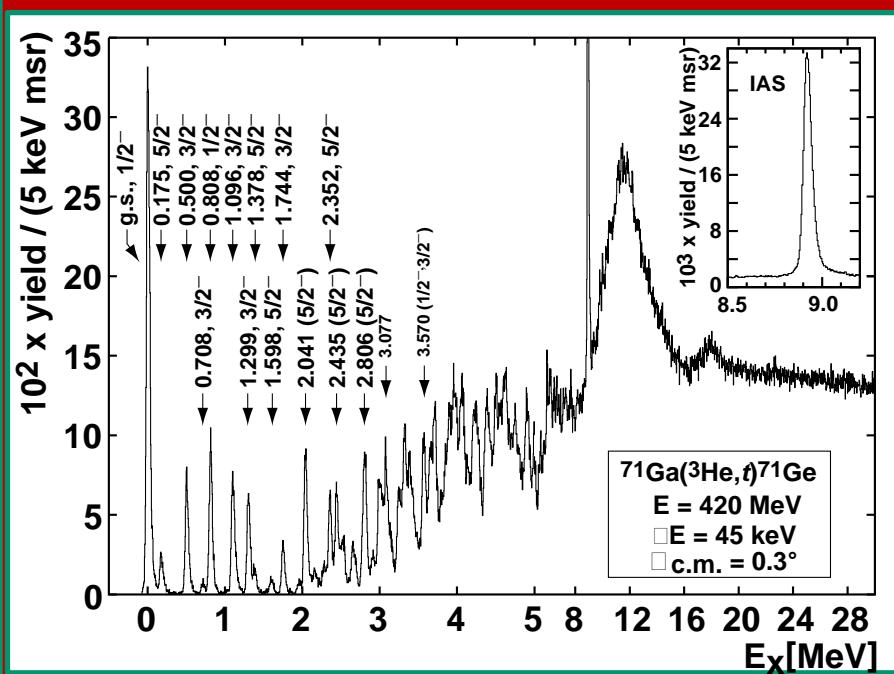




# solar neutrino rates via ( ${}^3\text{He}, t$ )

${}^{71}\text{Ga}(\nu_{\odot}, e^-)$  SNUs from  
 ${}^{71}\text{Ga}({}^3\text{He}, t){}^{71}\text{Ge}$  charge-ex reaction

# $^{71}\text{Ga}(\nu_{\odot}, e^-)$ SNUs from ( $^3\text{He}, t$ ) charge-exchange reaction



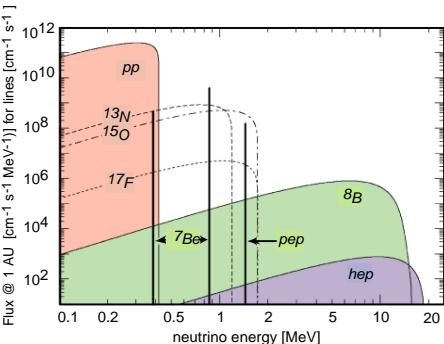
**71Ga( $\nu_{\odot}, e^-$ )**

$R = 122.4 \pm 3.4(\text{stat}) \pm 1.1(\text{sys})$

stat. err. mostly due to CNO ν's

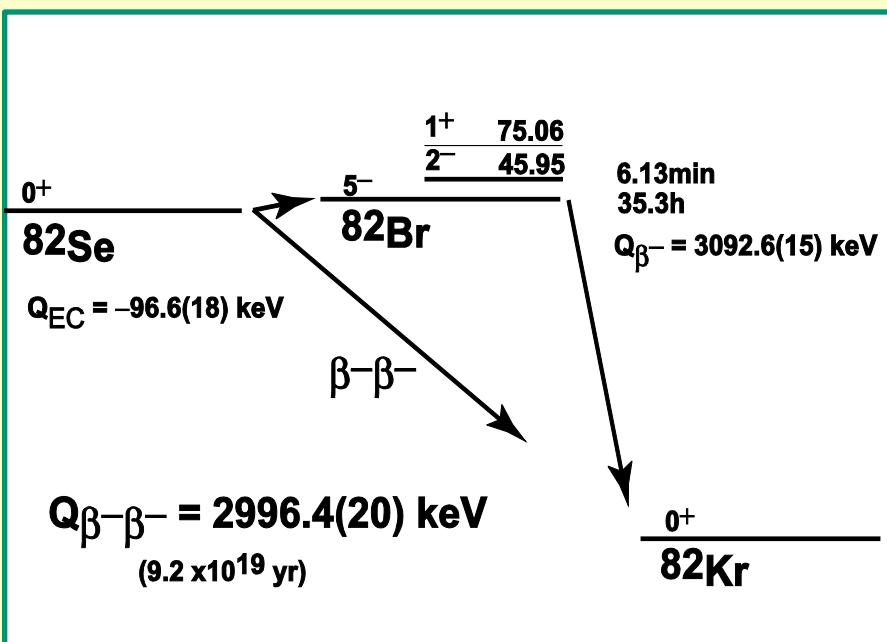
prev'ly:  $132 \pm 18$

DF et al, PRC91, 2015



# solar neutrino rates via ( ${}^3\text{He}, t$ )

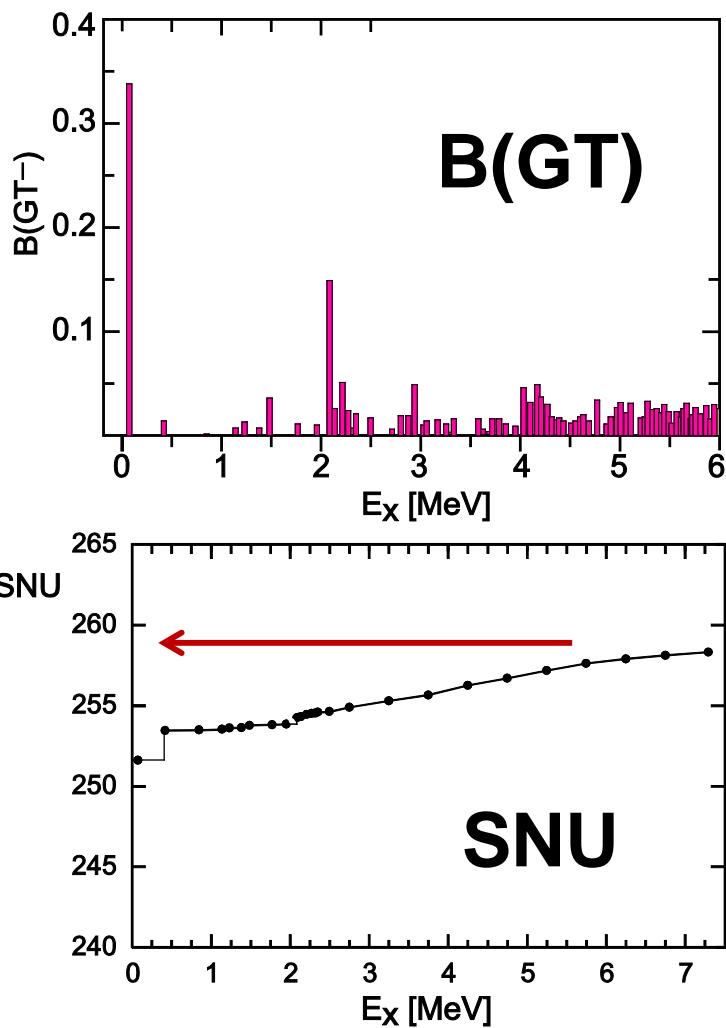
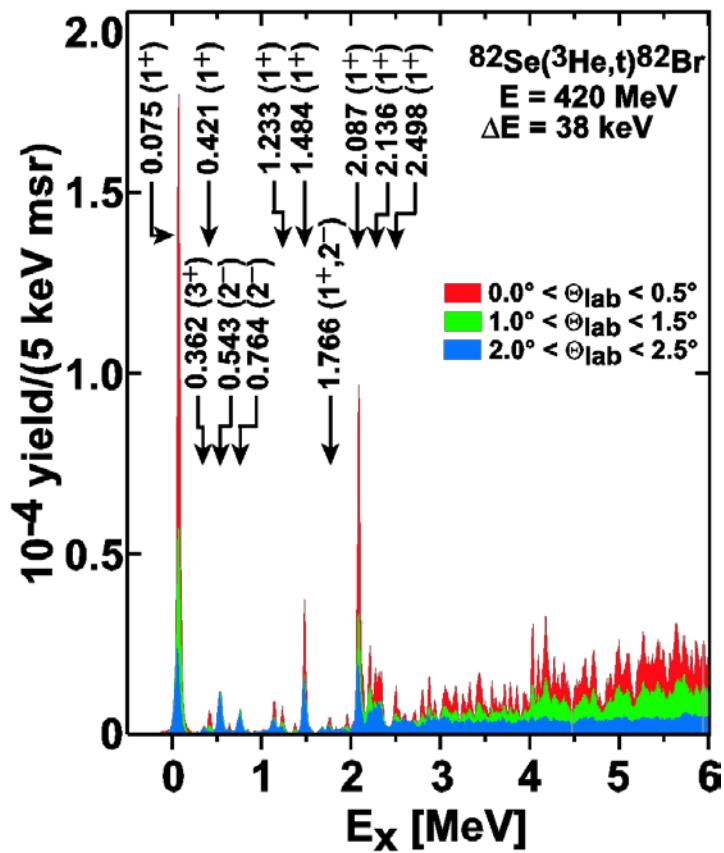
${}^{82}\text{Se}(\nu_{\odot}, e^-)$  SNUs from  
 ${}^{82}\text{Se}({}^3\text{He}, t){}^{82}\text{Br}$  charge-ex reaction



## Advantages:

- low threshold
- enhanced sensitivity to pp-neutrinos
- short life-time against  $\beta$ -decay (35h)
- pp- $\nu$ 's in „real time“
- $\gamma$ -emission, easy to detect

# $^{82}\text{Se}({}^3\text{He}, \text{t})$ spectrum



Total rate:

Population of 1<sup>st</sup> 1<sup>+</sup> state:

pp ν fraction:

258 SNU

97%

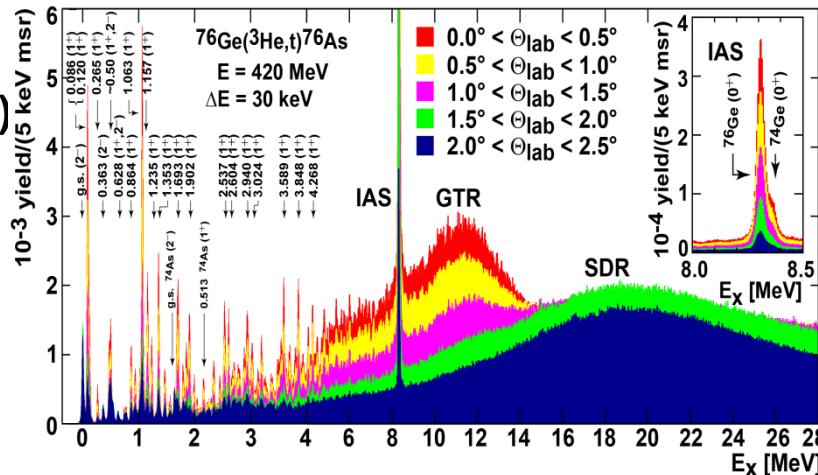
76%

preliminary

# Future perspectives of chargex-reactions

## ➤ $\beta\beta$ -decay and nuclear matrix elements

- Resolution is key issue (RCNP gives the lead!)
- need 20 - 30 keV for  $(^3\text{He}, t)$  &  $(d, ^2\text{He})$
- Need to explore proportionality between chargex x-section and  $\Delta L \neq 0$  transitions (e.g.  $2^-$  states) in weak interaction (resol'n is key)



## ➤ $\nu$ -physics and chargex-reactions

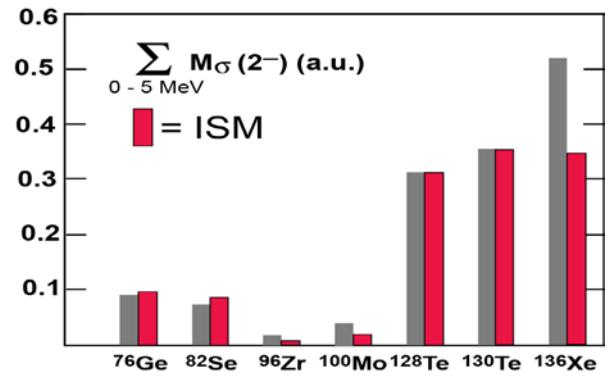
- Hadronic chargex and weak-interaction x-sections are fortuitously connected -- exploit this!!
- solar neutrinos, SN-neutrinos, element synthesis

## ➤ Need to address quenching issue urgently!!

- Chargex in inverse kinematics plays a pivotal role (BUT need resolution)

## ➤ EOS and chargex-reaction

- IAS and GT resonance data needed and useful BUT: theories need to converge on their relevance



$^{71}\text{Ga}(\nu_\odot, e^-)$

$$R = 122.4 \pm 3.4 \pm 1.1 \text{ SNU}$$

$^{82}\text{Se}(\nu_\odot, e^-)$

$$R = 258.4 \text{ SNU}$$

# Isobar separation method (U of Calgary)

