

Pairing fluctuations and giant dipole resonance

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Theoretical framework – hot nuclei





Microscopic-Macroscopic method

Strutinsky theorem :
$$F_{\text{TOT}} = E_{\text{LDM}} + \sum_{p,n} \delta F$$

 $\delta F = F - \tilde{F}$
 $F = E - TS = \sum_{i=1}^{\infty} e_i n_i - T \sum_{i=1}^{\infty} s_i$
 $\tilde{F} = 2 \sum_i e_i \tilde{n}_i^T - 2T \sum_i \tilde{s}_i + 2\gamma_s \int_{-\infty}^{\infty} \tilde{f}(x) x \sum_i n_i(x) dx$
At finite spin : $F_{\text{TOT}} = E_{\text{RLDM}} + \sum_{p,n} \delta F$
 $F_{\text{TOT}} = E_{\text{LDM}} + \sum_{p,n} \delta F^{\omega} + \frac{1}{2} \omega (I_{\text{TOT}} + \sum_{p,n} \delta I)$
 $\delta I = I - \tilde{I}$. $I_{\text{TOT}} = \Im_{rig} \omega + \delta I$

P. Arumugam, G. Shanmugam and S.K. Patra, PRC 69, 054313 (2004)
P. Arumugam, A.G. Deb and S.K. Patra, EPJA 25, 199 (2005)

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(b)

 $g(e) = \sum_{i=1}^{n} \frac{1}{4T \cosh^2[(e - e_i)/2T]}$

(a)



(c)

Model for GDR



Rotating anisotropic harmonic oscillator potential with a separable dipole-dipole residual interaction

 $H = H_{\rm av} + H_{\rm int}$

$$H_{\rm av}(\Omega) = \sum_{\nu=1}^{A} h_{\nu}(\Omega)$$

$$h(\Omega) = \frac{p^2}{2m} + \frac{m}{2} (\omega_x^2 x^2 + \omega_y^2 y^2 + \omega_z^2 z^2) - \Omega l_z \qquad L_z = \sum_{\nu=1}^A l_z(\nu)$$
$$H_{int} = \eta \sum_{i=x,y,z} \frac{m\omega_i^2}{2A} \left[\sum_{\nu=1}^A \tau_3^{(\nu)} x_i(\nu) \right]^2 \qquad \tau_3 = \begin{pmatrix} 1 & 0\\ 0 & -1 \end{pmatrix}$$

R.R. Hilton, **ZPA 309**, 233 (1983) G. Shanmugam and M.Thiagasundaram, **PRC 37**, 853 (1988); **PRC 39**, 1623 (1989)

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Model for GDR – contd.

In lab frame

 $\widetilde{\omega}_z = (1+\eta)^{1/2} \omega_z \; ,$

$$\begin{split} \tilde{\omega}_2 &\mp \Omega = \left\{ (1+\eta) \frac{\omega_y^2 + \omega_x^2}{2} + \Omega^2 + \frac{1}{2} \left[(1+\eta)^2 (\omega_y^2 - \omega_x^2)^2 \right. \\ &\left. + 8\Omega^2 (1+\eta) (\omega_y^2 + \omega_x^2) \right]^{\frac{1}{2}} \right\}^{\frac{1}{2}} \mp \Omega \;, \end{split}$$

$$\begin{split} \tilde{\omega}_3 &\mp \Omega = \begin{cases} (1+\eta) \frac{\omega_y^2 + \omega_x^2}{2} + \Omega^2 - \frac{1}{2} \left[(1+\eta)^2 (\omega_y^2 - \omega_x^2)^2 \right] \\ + 8\Omega^2 (1+\eta) (\omega_y^2 + \omega_x^2) \right]^{\frac{1}{2}} \end{split}^{\frac{1}{2}} \mp \Omega , \end{split}$$
 For spherical equations are consistent of the spherical equation in the spherical equation is the spherical equation of the spherical equation is the spherical equation equation equation equation equation equation equation

$$\sigma(E_{\gamma}) = \sum_{i} \frac{\sigma_{mi}}{1 + (E_{\gamma}^2 - E_{mi}^2)^2 / E_{\gamma}^2 \Gamma_i^2}$$
$$\sigma_m = 60 \frac{2NZ}{\pi A \Gamma} (1 + \alpha)$$

For collective rotation All 5 frequencies exist

For s.p. rotation
$$\widetilde{\omega}_1$$
 and $\widetilde{\omega}_2 - \Omega = \widetilde{\omega}_3 + \Omega$

ere $= \widetilde{\omega}_3 + \Omega$

 $\Gamma_i \approx 0.026 E_i^{1.9}$

 $\Gamma_i = \Gamma_0 (E_i / E_0)^{\delta}$





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Fig.1. Probability distributions for the gap parameter Δ at different termperatures. The value of Δ at the maximum corresponds to the solution of the gap equation. The critical temperature is T=0.57.



Fig.2. The average gap parameter (thick line) and the most probable gap parameter is a function of temperature.

Importance in GDR proposed by N.D. Dang, K. Tanabe, A. Arima, NPA675 (2000) 531

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Low-T GDR





¹⁷⁹Au at T = 0.7 MeV, Γ is quenched: F. Camera et al., PLB **560**, 155 (2003) Shell effects increase Γ: P. Arumugam, A.G. Deb and S.K. Patra, *Europhys. Lett.* **70**, 313 (2005)

TSFM with pairing: P. Arumugam, N. Dinh Dang, RIKEN Accel. Prog. Rep. 39 (2006) 28

Low-T GDR measurements at VECC, Kolkata PLB **709**, 9 (2012); PLB **713**, 434 (2012); PLB **731**, 92 (2014)



$$= H_{osc} + \eta D^{\dagger}D + \chi P^{\dagger}P$$

$$\omega_{\nu} = \omega_{\nu}^{OSC} - \chi \omega^{P} \qquad \qquad \omega^{P} = \left(\frac{Z\Delta_{P} + N\Delta_{N}}{Z + N}\right)^{2}$$

$$\eta = \eta_{0} - \chi_{0}\sqrt{T}\omega^{P}$$

H

$$\begin{split} \widetilde{\omega}_{1} &= (1+\eta)^{1/2} \,\omega_{z} ,\\ \widetilde{\omega}_{2} &= \left\{ (1+\eta) \frac{\omega_{y}^{2} + \omega_{x}^{2}}{2} + \frac{1}{2} \left\{ (1+\eta)^{2} (\omega_{y}^{2} - \omega_{x}^{2})^{2} \right\}^{1/2} \right\}^{1/2} \\ \widetilde{\omega}_{3} &= \left\{ (1+\eta) \frac{\omega_{y}^{2} + \omega_{x}^{2}}{2} - \frac{1}{2} \left\{ (1+\eta)^{2} (\omega_{y}^{2} - \omega_{x}^{2})^{2} \right\}^{1/2} \right\}^{1/2} \end{split}$$

Role of fluctuations







Free energy (F_{TOT}), calculated at abscissas using FTNS with exact T dependence Fluctuations in pairing field: λ is fixed at its BCS value and Δ is allowed to vary

$$\left\langle \mathcal{G} \right\rangle_{\beta,\gamma,\Delta_{P},\Delta_{N}} = \frac{\iint \int D[\alpha] e^{-F_{TOT}(T;\beta,\gamma,\Delta_{P},\Delta_{N})/T} \mathcal{G}}{\iint \int \int \int D[\alpha] e^{-F_{TOT}(T;\beta,\gamma,\Delta_{P},\Delta_{N})/T}}$$
$$D[\alpha] = \beta^{4} |\sin 3\gamma| \Delta_{P} \Delta_{N} d\beta d\gamma d\Delta_{P} d\Delta_{N}$$

A.K. Rhine Kumar, P. Arumugam and N.D. Dang, PRC 91, 044305 (2015)



Role of pairing in ¹²⁰Sn



A.K. Rhine Kumar, P. Arumugam and N.D. Dang, PRC 91, 044305 (2015); PRC 90, 044308 (2014).

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Testing prescriptions for G

G from Strutinsky calculations



G from empirical data

$$G_{P,N} = [19.2 \pm 7.4(N-Z)]/A^2$$

P. Arumugam, N. Dinh Dang, RIKEN Accel. Prog. Rep. **39** (2006) 28









 $\text{PF}(\text{GCE})^*$ correspond to adjusted energy dependence of Γ



• Precisely measured shift in E_{GDR} could reveal more about χ

- Only a qualitative fit is possible
- Analysis with BW form of cross sections is in progress

Expt. Data: PLB **731**, 92 (2014)



Nucleus	(i) Shell	(ii) Pairing effects through $F_{\rm TOT}$			(iii)Pairing effects	Net
	effects	$T = 0.1 { m ~MeV}$	$T=0.4~{\rm MeV}$	$T = 0.6 { m ~MeV}$	through $\chi \mathcal{P}^{\dagger} \mathcal{P}$	effect
$^{120}\mathrm{Sn}$		1	\downarrow	↓	\downarrow	\downarrow
$^{179}\mathrm{Au}$	1	Ļ	\downarrow	\downarrow	\downarrow	\downarrow
$^{208}\mathrm{Pb}$	\downarrow				\downarrow	\downarrow
$^{97}\mathrm{Tc}$	1	↓	\downarrow		\downarrow	↓

Hot and rotating ¹⁵²Gd

 $\Gamma_{\rm D}$ (MeV)



¹⁵²Gd (²⁸Si+¹²⁴Sn) at beam energy 185 MeV



Tuning parameters to fit Γ could be misleading. Comparison with cross-sections is more robust.

A.K. Rhine Kumar and P. Arumugam, to appear in PRC

¹⁵²Gd at beam energy 149 MeV





A.K. Rhine Kumar and P. Arumugam, to appear in PRC

Summary



- Thermal fluctuations in pairing (PF) leads to a quenching of GDR width in agreement with the experimental observations.
- For small systems like the atomic nuclei, there are no sharp superfluidnormal phase transitions.
- For a precise match with the ¹²⁰Sn and ⁹⁷Tc data, the consideration of PF is crucial.
- In ²⁰⁸Pb, the shell effects are so dominating as to favour a spherical (closed-shell/unpaired) configuration and hence the role of PF is negligible.
- One may consider the precise low-*T* GDR measurements as a benchmark for pairing prescriptions.
- ¹⁵²Gd: Tuning parameters to fit Γcould be misleading. Comparison with cross-sections is more robust.

Thank You



Energy dependence of width

- This is a profound question and most of the answers are tentative
 The feasible approach with predictive power is only empirical
 - Г (MeV) $\Gamma_i \approx 0.026 E_i^{1.9}$ P. Carlos et al, NPA219, 61 (1974) 208 4 Only few nuclei were chosen Э 15 E (MeV) 12 13 14 Fig. 9. Experimental Γ_{i}^{\downarrow} values (•) and best fit (dotted line) corresponding to $\Gamma^{\downarrow} = 0.026 E^{1.9}$.







Woods-Saxon Strutinsky results





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